



Conformal symmetry of inflationary perturbations

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+ work in progress with M. Simonovic and J. Noreña
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Test scalars in de Sitter have conformal-invariant correlation functions

The consistency relation for the 3-point function does not receive linear corrections

$$\langle \zeta(\vec{k}_1) \zeta(\vec{k}_2) \zeta(\vec{k}_3) \rangle \simeq -(2\pi)^3 \delta(\vec{k}_1 + \vec{k}_2 + \vec{k}_3) P(k_1) P(k_S) \left[\frac{d \ln(k_S^3 P(k_S))}{d \ln k_S} + \mathcal{O}\left(\frac{k_1^2}{k_S^2}\right) \right]$$

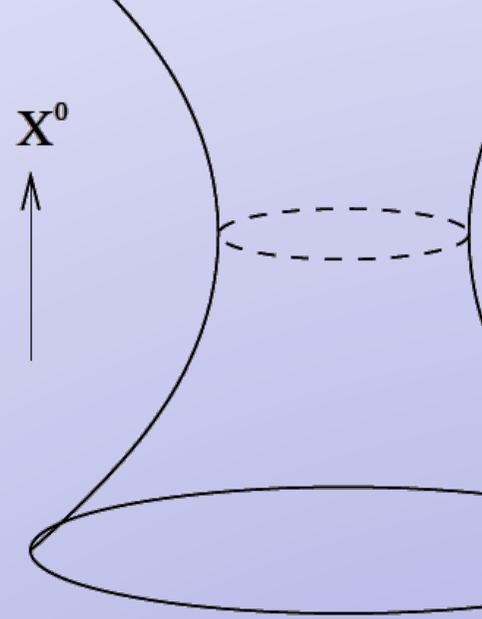
$k_1 \rightarrow 0$

Non-linear realization of conformal invariance constrains the squeezed limit of correlation functions

New consistency relations

Inflation takes place in $\sim dS$

$$ds^2 = \frac{1}{H^2 \eta^2} (-d\eta^2 + d\vec{x}^2)$$



Translations, rotations: ok

Dilations (if slow-roll)

$$\eta \rightarrow \lambda \eta, \quad \vec{x} \rightarrow \lambda \vec{x}$$

→ scale-invariance

$$\varphi_{\vec{k}} \rightarrow \lambda^3 \varphi_{\lambda \vec{k}}$$

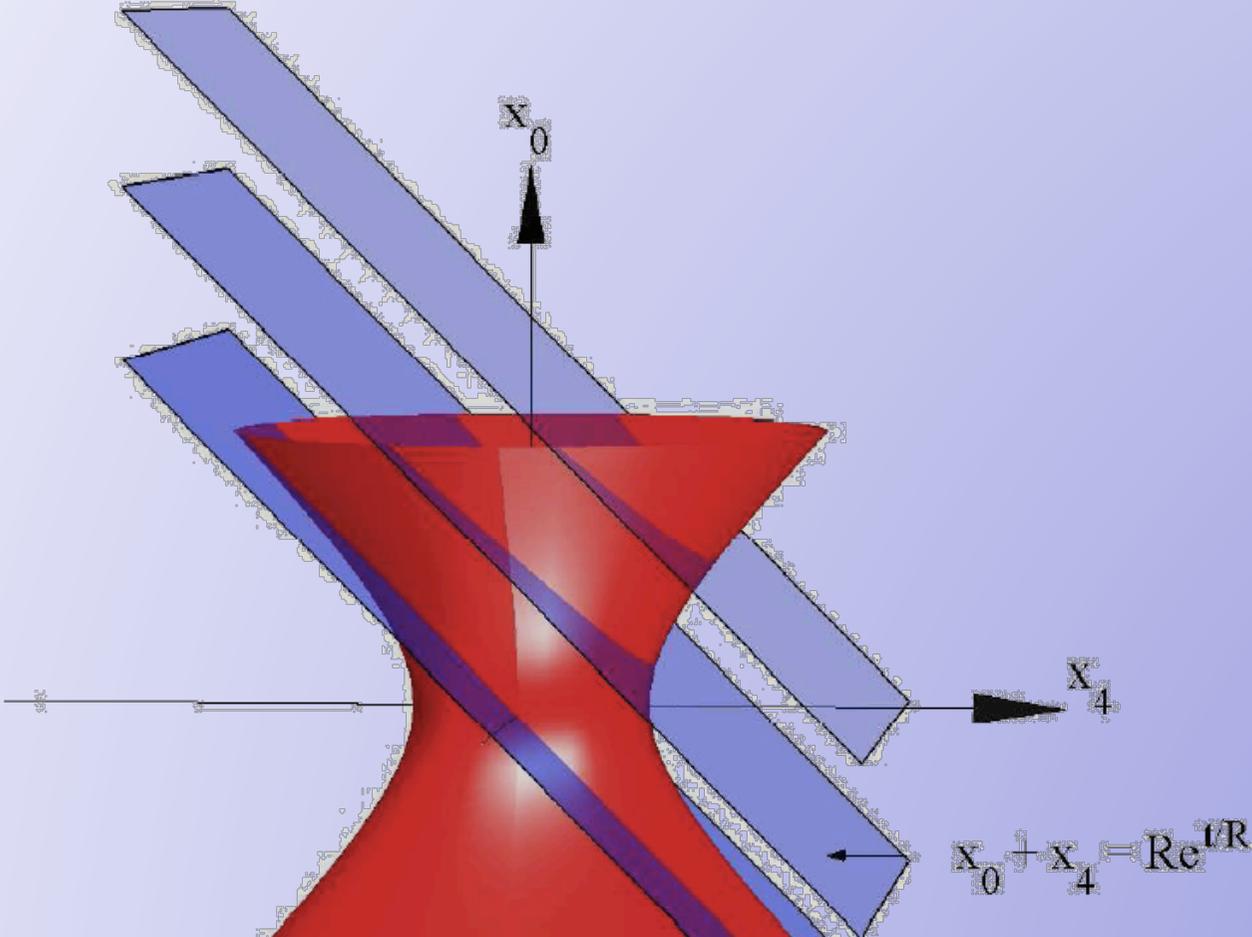
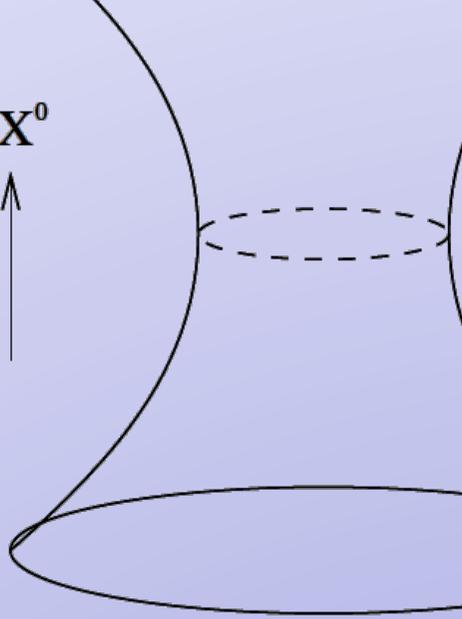
$$\langle \varphi_{\vec{k}_1} \varphi_{\vec{k}_2} \rangle = (2\pi)^3 \delta(\vec{k}_1 + \vec{k}_2) \frac{1}{k_1^3} F(k_1 \eta)$$

Special conformal

$$\eta \rightarrow \eta - 2\eta(\vec{b} \cdot \vec{x}), \quad x^i \rightarrow x^i + b^i(-\eta^2 + \vec{x}^2) - 2x^i(\vec{b} \cdot \vec{x})$$

Inflation takes place in $\sim dS$

$$ds^2 = \frac{1}{H^2 \eta^2} (-d\eta^2 + d\vec{x}^2)$$



perturbations are created by a sector with negligible interactions with the inflator
correlation functions have the full $SO(4,1)$ symmetry

They are conformal invariant

Independently of any details about this sector, even at strong coupling

We are interested in correlators at late times

$$x^i \rightarrow x^i + b^i \vec{x}^2 - 2x^i (\vec{b} \cdot \vec{x}) \quad \eta \rightarrow \eta - 2\eta(\vec{b} \cdot \vec{x})$$

$$\varphi \sim \eta^\Delta, \quad \Delta = \frac{3}{2} \left(1 - \sqrt{1 - \frac{4m^2}{9H^2}} \right) \ll 1$$

This is the transformation of the a primary of conformal dim Δ

ample: $m = \sqrt{2}H \quad \Delta = 1$

$$d^4x \sqrt{-g} \frac{M}{6} \varphi^3 \quad \longrightarrow \quad \langle \varphi_{\vec{k}_1} \varphi_{\vec{k}_2} \varphi_{\vec{k}_3} \rangle = (2\pi)^3 \delta\left(\sum_i \vec{k}_i\right) \frac{\pi}{16} M H^2 \eta_*^3 \cdot \frac{1}{k_1 k_2 k_3}$$

$$\langle \varphi(\vec{x}_1) \varphi(\vec{x}_2) \varphi(\vec{x}_3) \rangle = \frac{M H^2 \eta_*^3}{128 \pi^2} \cdot \frac{1}{|\vec{x}_1 - \vec{x}_2| |\vec{x}_1 - \vec{x}_3| |\vec{x}_2 - \vec{x}_3|}$$

$$\langle \varphi_{\vec{k}_3} \rangle = (2\pi)^3 \delta(\sum_i \vec{k}_i) \frac{H^2}{\prod_i 2k_i^3} \frac{2M}{3} \left[\sum_i k_i^3 (-1 + \gamma + \log(-k_t \eta_*)) + k_1 k_2 k_3 - \sum_i \dots \right]$$

$$k_t \equiv \sum_i k_i$$

Zaldarriaga '0
Seery, Malik, Lyt

$$\langle \varphi_1(\vec{x}_1) \varphi_2(\vec{x}_2) \varphi_3(\vec{x}_3) \rangle = \frac{MH^2}{48\pi^2} \log \frac{|\vec{x}_1 - \vec{x}_2|}{A\eta_*} \log \frac{|\vec{x}_1 - \vec{x}_3|}{A\eta_*} \log \frac{|\vec{x}_2 - \vec{x}_3|}{A\eta_*}$$

Everything determined up to two constants

Independently of the interactions!

$$\int d^4x \sqrt{-g} \nabla_\mu \varphi_1 \nabla^\mu \varphi_2 \varphi_3 \longrightarrow \frac{1}{M} \int d^4x \sqrt{-g} \frac{1}{2} (\square \varphi_3 \varphi_1 \varphi_2 - \square \varphi_1 \varphi_2 \varphi_3 - \square \varphi_1 \varphi_3 \varphi_2)$$

$$\int d^4x \frac{1}{8M^4} (\partial_\mu \varphi)^2 (\partial_\nu \varphi)^2$$

$$\langle \varphi_{\vec{k}_1} \varphi_{\vec{k}_2} \varphi_{\vec{k}_3} \varphi_{\vec{k}_4} \rangle = (2\pi)^3 \delta(\sum_i \vec{k}_i) \frac{1}{M^4} \frac{H^8}{\prod_i 2k_i^3} \left[-\frac{144k_1^2 k_2^2 k_3^2 k_4^2}{k_t^5} - 4 \left(\frac{12k_1 k_2 k_3 k_4}{k_t^5} + \frac{3 \prod_{i<j<l} k_i k_j k_l}{k_t^4} + \frac{\prod_{i<j} k_i k_j}{k_t^3} \right) \right. \\ \left. + (\vec{k}_3 \cdot \vec{k}_4) + (\vec{k}_1 \cdot \vec{k}_3)(\vec{k}_2 \cdot \vec{k}_4) + (\vec{k}_1 \cdot \vec{k}_4)(\vec{k}_2 \cdot \vec{k}_3) + (\vec{k}_1 \cdot \vec{k}_2) \left(\frac{4k_3^2 k_4^2}{k_t^3} + \frac{12(k_1 + k_2)k_3^2 k_4^2}{k_t^4} + \frac{48k_1 k_2 k_3^2 k_4^2}{k_t^5} \right) + 6 \text{perm} \right]$$

Not so obvious it is conformal invariant...

I can check it in **Fourier space**

Maldacena and Pirm

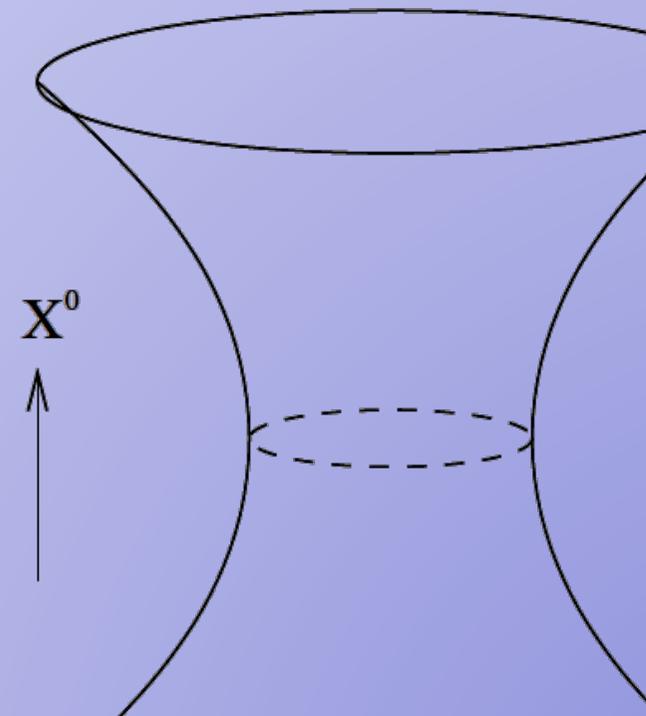
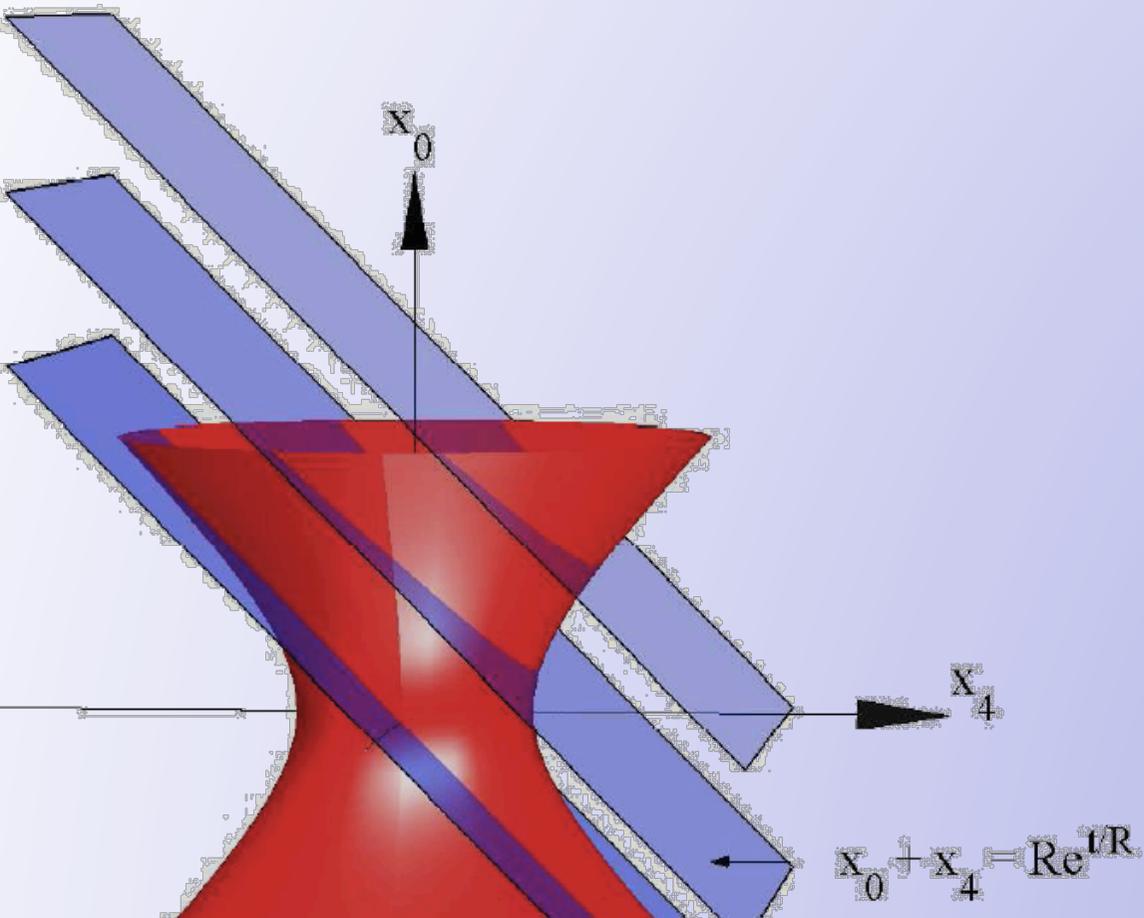
$$\sum_{a=1,2,3,4} \left[6\vec{b} \cdot \vec{\partial}_{k_a} - \vec{b} \cdot \vec{k}_a \vec{\partial}_{k_a}^2 + 2\vec{k}_i \cdot \vec{\partial}_{k_a} (\vec{b} \cdot \vec{\partial}_{k_a}) \right] \langle \varphi_{\vec{k}_1} \varphi_{\vec{k}_2} \varphi_{\vec{k}_3} \varphi_{\vec{k}_4} \rangle' = 0$$

general: $F \left(\frac{r_{13} r_{24}}{r_{12} r_{34}}, \frac{r_{23} r_{41}}{r_{12} r_{34}} \right) \prod_{i<j} r_{ij}^{-2\Delta/3}$

2 parameters instead of 5

If we see something beyond the spectrum

- Something not conformal would be a probe of a "sliced" de Sitter
- Something conformal would be a probe of pure de Sitter



Squeezed limit in single-field models: one of the modes is a classical bkg when the other two exit the H^{-1}

$$\langle \zeta_B(\vec{k}_1) \zeta(\vec{k}_2) \zeta(\vec{k}_3) \rangle \simeq \langle \zeta_B(\vec{k}_1) \langle \zeta(\vec{k}_2) \zeta(\vec{k}_3) \rangle_B \rangle \quad k_1 \ll k_S$$

The long mode acts just as a rescaling of the coordinates

$$\langle \zeta(\vec{x}_2) \zeta(\vec{x}_3) \rangle_B = \xi(\vec{x}_2 - \vec{x}_3) \simeq \xi(\vec{x}_3 - \vec{x}_2) + \zeta_B(\vec{x}_+) (\vec{x}_3 - \vec{x}_2) \cdot \nabla \xi(\vec{x}_3 - \vec{x}_2)$$

Going back to Fourier space we get the consistency relation

$$\langle \zeta(\vec{k}_1) \zeta(\vec{k}_2) \zeta(\vec{k}_3) \rangle \simeq -(2\pi)^3 \delta(\vec{k}_1 + \vec{k}_2 + \vec{k}_3) P(k_1) P(k_S) \frac{d \ln(k_S^3 P(k_S))}{d \ln k_S}$$

$$\vec{k}_+ = (\vec{k}_1 + \vec{k}_2) / 2$$

lowest order in derivatives

$$S_2 + S_3 = M_{\text{Pl}}^2 \int d^4x \epsilon a^3 \left[(1 + 3\zeta_B) \dot{\zeta}^2 - (1 + \zeta_B) \frac{(\partial_i \zeta)^2}{a^2} \right]$$

g mode reabsorbed by coordinate rescaling $\vec{x} \rightarrow (1 + \zeta_B) \vec{x}$

rections:

me evolution of ζ is of order k^2

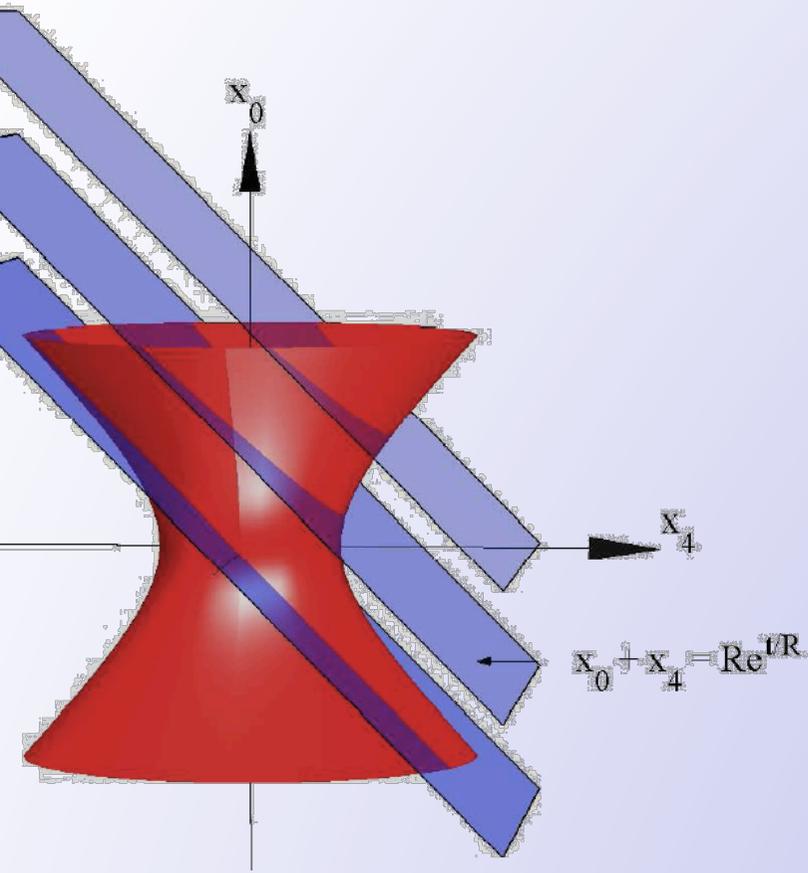
patial derivatives will be symmetrized with the short modes, giving k^2

onstraint equations give order k^2 corrections

Final result: in the not-so-squeezed limit we have

$$\langle \zeta(\vec{k}_1) \zeta(\vec{k}_2) \zeta(\vec{k}_3) \rangle \simeq -(2\pi)^3 \delta(\vec{k}_1 + \vec{k}_2 + \vec{k}_3) P(k_1) P(k_S) \left[\frac{d \ln(k_S^3 P(k_S))}{d \ln k_S} + \mathcal{O}\left(\frac{k_1^2}{k_S^2}\right) \right]$$

The inflaton does not break $SO(4,1)$ explicitly: it is still non-linearly realized



$$\phi(\vec{x}, t) = t + \pi(\vec{x}, t)$$

$$\eta \rightarrow \eta - 2\eta(\vec{b} \cdot \vec{x})$$

$$x^i \rightarrow x^i + b^i(-\eta^2 + \vec{x}^2) - 2x^i(\vec{b} \cdot \vec{x})$$

$$\pi \rightarrow \pi + 2H^{-1}(\vec{b} \cdot \vec{x})$$

A gradient of a background mode can be traded for a conformal transformation



(Assuming zero tilt for simplicity)

$$\langle \zeta_{\vec{q}} \zeta_{\vec{k}_1} \zeta_{\vec{k}_2} \zeta_{\vec{k}_3} \dots \rangle' = -\frac{1}{2} P(q) q^i D_i \langle \zeta_{\vec{k}_1} \zeta_{\vec{k}_2} \zeta_{\vec{k}_3} \dots \rangle', \quad |\vec{q}| \rightarrow 0$$

$$q^i D_i = \sum_{a=1}^n \left[6\vec{q} \cdot \vec{\partial}_{k_a} - \vec{q} \cdot \vec{k}_a \vec{\partial}_{k_a}^2 + 2\vec{k}_a \cdot \vec{\partial}_{k_a} (\vec{q} \cdot \vec{\partial}_{k_a}) \right]$$

- The variation of the 2-point function is zero: no linear term in the 3 pf

Linear terms in the 4pf are connected with the conf variation of the 3-point function

- No tilt suppression

It encodes at the level of observables the relation among operators

$$(g^{00} + 1)^2 \supset -2\dot{\pi}(\partial_{\mu}\pi)^2 + (\partial_{\mu}\pi)^4$$

- ✓ $\dot{\pi}^4$ has free coefficient, indeed it has no k^{-2} squeezed limit
- ✓ $(\nabla\pi)^4$ and $\dot{\pi}^2(\nabla\pi)^2$ are related to $\dot{\pi}(\nabla\pi)^2$ and $\dot{\pi}^3$

Also including gravity

So far we have neglected metric perturbations + taken exact dS but everything still holds

$$x^i \rightarrow x^i + b^i \vec{x}^2 - 2x^i (\vec{b} \cdot \vec{x})$$

$$\langle \zeta_{\vec{k}_1} \zeta_{\vec{k}_2} \zeta_{\vec{k}_3} \rangle = (2\pi)^3 \delta^3 \left(\sum \vec{k}_i \right) \frac{\dot{\rho}_*^4 H_*^4}{\dot{\phi}_*^4 M_{pl}^4} \frac{1}{\prod_i (2k_i^3)} \mathcal{A}_*$$

$$\mathcal{A} = 2 \frac{\ddot{\phi}_*}{\dot{\phi}_* \dot{\rho}_*} \sum_i k_i^3 + \frac{\dot{\phi}_*^2}{\dot{\rho}_*^2} \left[\frac{1}{2} \sum_i k_i^3 + \frac{1}{2} \sum_{i \neq j} k_i k_j^2 + 4 \frac{\sum_i k_i^2 k_j^2}{\dots} \right]$$

$$\langle \zeta_{\mathbf{k}_1} \zeta_{\mathbf{k}_2} \zeta_{\mathbf{k}_3} \zeta_{\mathbf{k}_4} \rangle^{\text{CI}} = (2\pi)^3 \delta \left(\sum_a \mathbf{k}_a \right) \frac{H_*^6}{4\epsilon^2 \prod_a (2k_a^3)} \sum_{\text{perms}} \mathcal{M}_4(\mathbf{k}_1, \mathbf{k}_2, \mathbf{k}_3, \mathbf{k}_4)$$

$$\begin{aligned} \mathcal{M}_4(\mathbf{k}_1, \mathbf{k}_2, \mathbf{k}_3, \mathbf{k}_4) = & -2 \frac{k_1^2 k_3^2}{k_{12}^2 k_{34}^2} \frac{W_{24}}{k_t} \left(\frac{\mathbf{Z}_{12} \cdot \mathbf{Z}_{34}}{k_{34}^2} + 2\mathbf{k}_2 \cdot \mathbf{Z}_{34} + \frac{3}{4} \sigma_{12} \sigma_{34} \right) \\ & - \frac{1}{2} \frac{k_3^2}{k_{34}^2} \sigma_{34} \left(\frac{\mathbf{k}_1 \cdot \mathbf{k}_2}{k_t} W_{124} + 2 \frac{k_1^2 k_2^2}{k_t^3} + 6 \frac{k_1^2 k_2^2 k_4}{k_t^4} \right), \end{aligned}$$

$$\sigma_{ab} = \mathbf{k}_a \cdot \mathbf{k}_b + k_b^2,$$

$$\mathbf{Z}_{ab} = \sigma_{ab} \mathbf{k}_a - \sigma_{ba} \mathbf{k}_b,$$

$$W_{ab} = 1 + \frac{k_a + k_b}{k_t} + \frac{2k_a k_b}{k_t^2},$$

- Linear and non-linear realization of conformal invariance
- Future directions:
 - Relation with Ward identities for spontaneously broken symmetries
 - Extension to models with $SO(4,2)$
 - Extension to correlators involving gravitons