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# Calorimetric approach to the direct measurement of the neutrino mass: the MARE project

Erica Andreotti

University of Insubria & INFN – Milano Bicocca

on behalf of the MARE collaboration



# Summary

- The physical context
- Basic concepts & experimental requirements for the direct measurement of  $m_\nu$
- Spectrometers versus Calorimeters
- Re-based  $\mu$ calorimeters: basic concepts & state-of-the-art (special care to MIBETA & semiconductor thermistors)
- The MARE project: aims, potentiality & experimental requirements

# The physical context

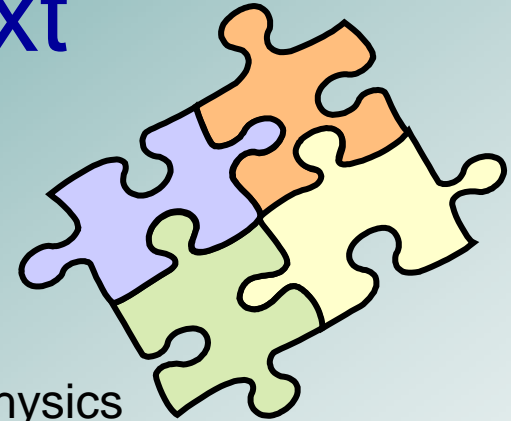


Oscillation experiments



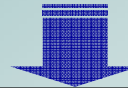
Neutrinos have non-zero mass

$\nu$  mass scale is crucial over 2 fronts:  
elementary particles physics & astroparticle physics



WE NEED TO KNOW:

- absolute  $\nu$  mass scale
- nature of  $\nu$  mass



Direct search  
through  $\beta$  decay

Potential sensitivity  
 $m(\nu_e) \sim 0.2 \text{ eV}/c^2$

- ❖ in the game if  $m_\nu$  quasi-degenerate
- ❖ completely model free

Cosmology

$$\sum m_{\nu_i} \leq 1 \text{ eV}/c^2$$

- ❖ planned sensitivity 0.05 eV
- ❖ spread in recent results
- ❖ model dependent

$0\nu\text{-}\beta\beta$  decay

$$\langle m_\nu \rangle = 0.4 \text{ eV}/c^2$$

(to be confirmed)

- ❖ planned sensitivity 0.05 eV
- ❖ works only if neutrino is a Majorana particle

# Direct $m_\nu$ measurements via $\beta$ decay



$$\longrightarrow E_0 = M_{\text{at}}(A,Z) - M_{\text{at}}(A,Z+1) \cong E_e + E_\nu$$

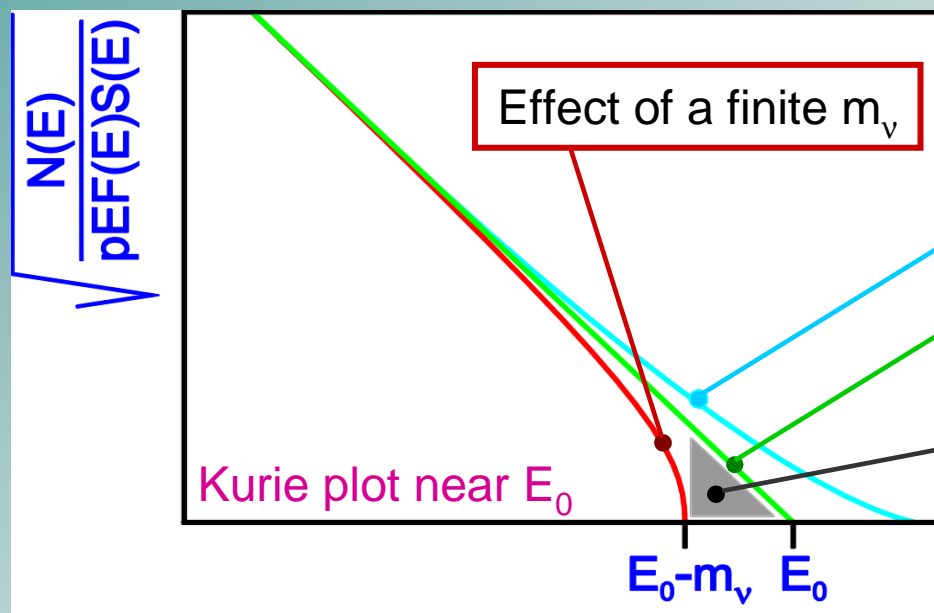
basic idea:

study the distribution of the  $e^-$  energy in proximity of the end-point

For a **finite**  $m_\nu$ :

$$N_\beta(Z, E_\beta, m_{\nu_e}) = p_\beta E_\beta (E_0 - E_\beta) \sqrt{(E_0 - E_\beta)^2 - m_{\nu_e}^2 c^4} F(Z, E_\beta) S(E_\beta) [1 + \delta_R(Z, E_\beta)]$$

Kurie plot  $K(E_e)$ : convenient **linearization** of the beta spectrum



- effect of:
- background
  - energy resolution
  - excited final states
  - pile-up

Fraction of decays below endpoint:

$$F(\delta E) = \int_{E_0 - \delta E}^{E_0} N(E_\beta, m_\nu=0) dE \cong 2 \left[ \frac{\delta E}{E_0} \right]^3$$

# Experimental requirements

- High statistics at the end-point

$$F(\delta E) \sim 2 \left[ \frac{\delta E}{E_0} \right]^3 \quad \Rightarrow \quad \text{Low } E_0 \beta\text{-decaying isotopes required!}$$

- High energy resolution  $\Rightarrow$  a tiny spectral distortion must be observed

- Approximate evaluation of sensitivity to  $m_\nu$

$$\sigma(M_\nu) \cong \sqrt[4]{\frac{1.6 E_0^3 \Delta E}{A T_M}}$$

High energy resolution (points to  $\Delta E$ )  
 total source activity (points to  $A$ )  
 live time (points to  $T_M$ )  
 High statistics (bracketed around the denominator terms)

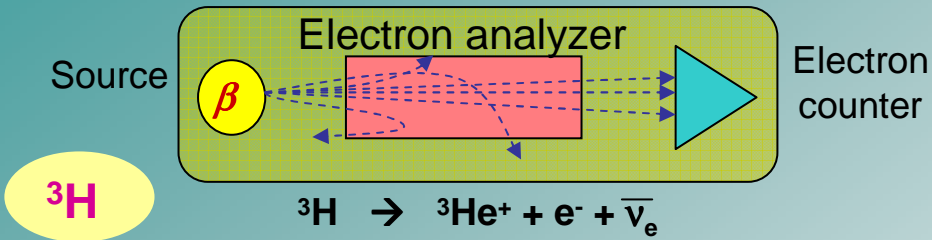
- Small & well known systematic effects  $\Rightarrow$  they could distort the spectral shape
  - unaccounted background gives negative  $m(\nu_e)^2$
  - response of the detector (i.e. energy resolution)
  - problem of excited final state
  - pile-up effects

**➔ At least two different & complementary approaches required!**

# Two complementary experimental approaches

spectrometer  
source ≠ detector

Only the useful fraction of electrons with  
 $E \sim E_0$  selected



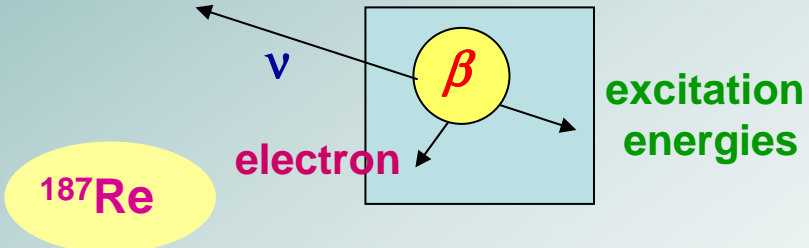
$E_0 = 18.57 \text{ keV}$  → high statistics at the end-point  
 $T_{1/2} = 12.3 \text{ y}$  → high specific activity  $\beta$ -source  
 Superallowed transition → no problem for analytical determination of  $\beta$ -spectrum

- Many sources of systematics:**
- deconvolve **detector response function**
  - **self-absorption** in the source
  - **inelastic scattered** electrons
  - problem of **final excited states**

**90's MAINZ-TROITZK** ⇒  $m_\nu < 2.2 \text{ eV}$   
**KATRIN** will start in 2011 ⇒  $m_\nu \sim 0.2 \text{ eV}$

calorimeter  
source = detector

Determines all the visible energy:  
→  $\nu$  energy as a **missing energy**



$E_0 = 2.47 \text{ keV}$  → the lowest in nature!  
 $T_{1/2} = 43.2 \text{ Gy}$  → 1 Bq/mg, ideal for bolometers  
 Unique 1<sup>st</sup> forbidden transition → computable nuclear matrix element

- Main advantage: excited final states
- Main problem: pile-up

**Completely different systematics!**

**MIBETA, MANU** ⇒  $m_\nu < 15.0 \text{ eV}$   
**Future: MARE** ⇒  $m_\nu \sim 0.2 \text{ eV}$



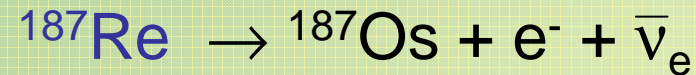
# Re-based $\mu$ calorimeters

Calorimeters measure the **entire spectrum** at once

→ low  $E_0$  to achieve enough statistic close to  $E_0$

Best choice:  $^{187}\text{Re} > E_0 = 2.47 \text{ keV}$

event frac. in the last 10 eV:  $1.3 \times 10^{-7}$  vs.  $3 \times 10^{-10}$  for  $^3\text{H}$  beta spectrum



Large isotopic abundance: 62.8%

No need of isotopic separation

• Main advantage: excited final states

The neutrino energy is measured as a “**missing energy**”

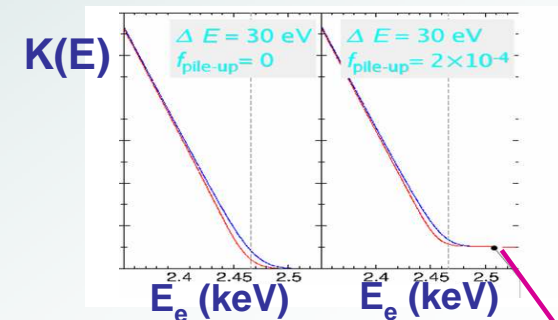
When in presence of decays to **excited states**, the calorimeter measures the **de-excitation energy**

• Main problem: pile-up

Bolometers **intrinsically slow** + **whole**  $\beta$  spectrum acquired

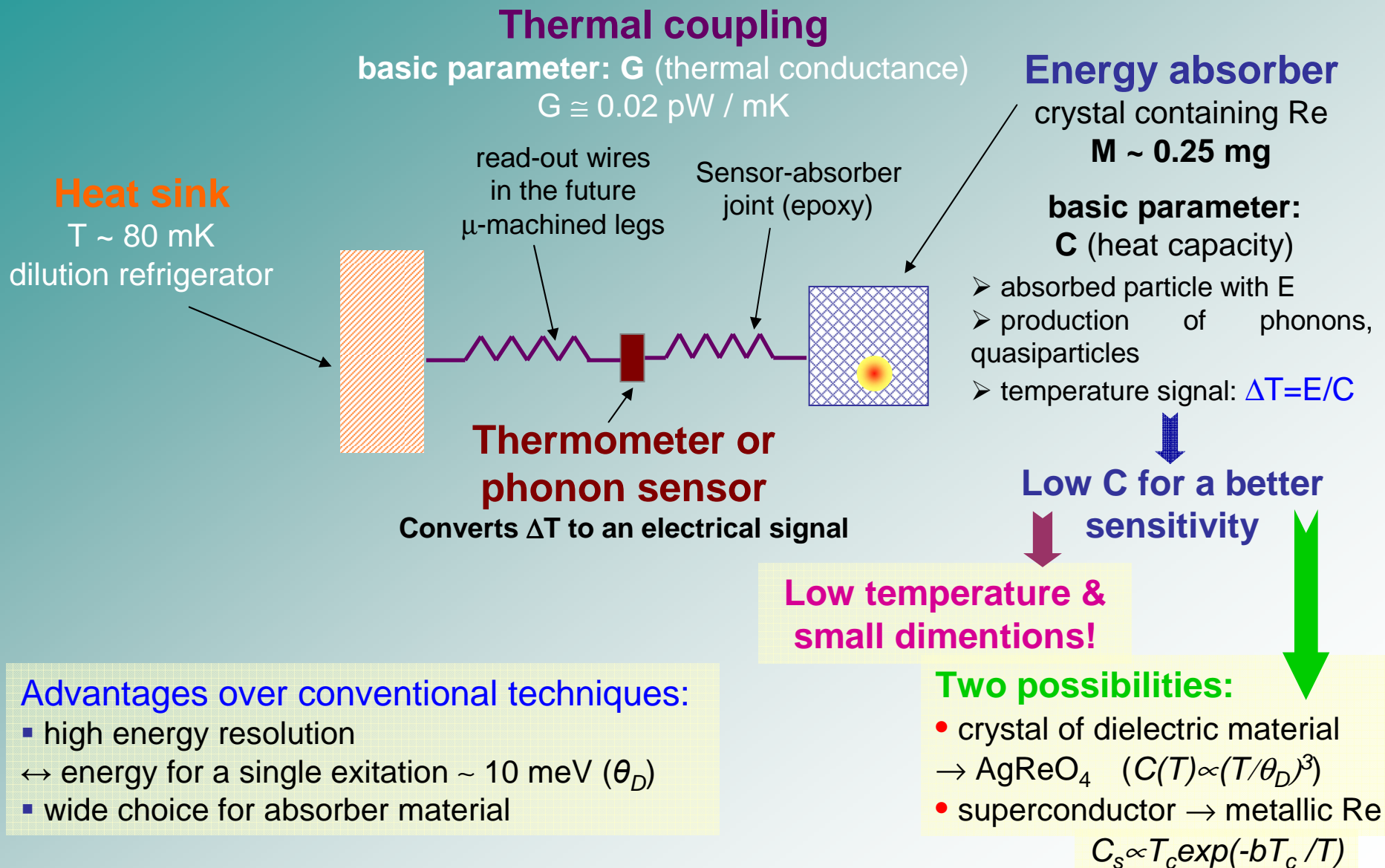
$^{187}\text{Re} - T_{1/2} = 43.2 \text{ Gy}$   
→ 1Bq/mg: ideal!

Development of arrays of many small detectors ( **$\mu$ calorimeters**)  
High reproducibility required!



pile-up fraction  $\sim A \times \tau_r$

# Bolometric detectors of particles: basic concepts



## Advantages over conventional techniques:

- high energy resolution  
 $\leftrightarrow$  energy for a single excitation  $\sim 10 \text{ meV}$  ( $\theta_D$ )
- wide choice for absorber material



# Thermistors technology & precursors experiments

**MIBETA**  
 Milan/Como 2000-03  
 $\text{AgReO}_4$   
 $\langle M_v \rangle < 15 \text{ eV}$  (90% C.L.)

**MANU**  
 Genoa 1995-99  
 Metallic Re  
 $\langle M_v \rangle < 19 \text{ eV}$  (90% C.L.)

## Thermometer:

Resistive element with heavy dependence of the resistance on the temperature  
**CRITICAL PARAMETERS:**  $\tau_r$  and Signal/Noise

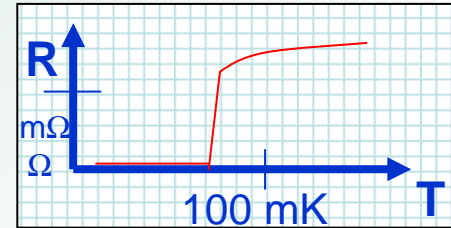
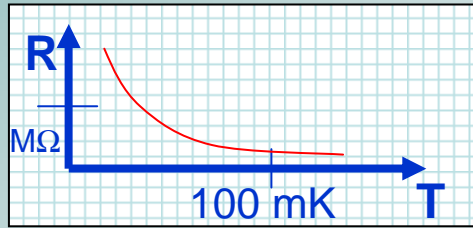
Specific know-how developed on semiconductor thermistor technology

Specific know-how developed on transition edge sensors (TES)

**Wisconsin-NASA**  
 $\mu$ calorimeter arrays for application to X-ray astronomy

**Variable Range Hopping (VRH)**  
 conduction regime: **exponential** increase of R with decreasing T

**Superconducting film** ( $\sim 10^2 \text{ nm}$ ) deposited on the absorber kept at the transition edge  $T_c \rightarrow$  resistivity changes rapidly with temperature fluctuation



# Semiconductor thermistors

**Ge** or **Si** small crystals doped slightly below the **metal-insulator transition** (MIT)

**Variable Range Hopping**  
conduction regime:  
phonon assisted tunneling of  $e^-$

$$R(T) = R_0 \exp(T_0/T)^p$$

$T < 1K$   $p=1/2$   
 $T_0$  &  $R_0$  depends on doping level

High  $T_0$

Accurate characterization of  
thermistors to optimize  $S/N$  &  $\tau_r$

Low  $T_0$  + large  
volume

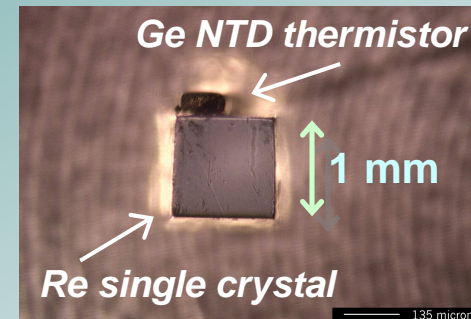
Small volume

Other factors  
(electron-phonon decoupling)

**C** must not exceed the  
absorber heat capacity

# Precursors experiments: MANU (Genoa 1995-99)

- **Metallic** Re single crystal → **ONE** detector only
- mass 1.6 mg → Activity  $A_\beta = 1.6$  Hz
- thermometer: **Ge NTD thermistor** (VRH),  
size = 0.1 x 0.1 x 0.23 mm<sup>3</sup>
- live time: 0.5 year
- $\Delta E_{\text{FWHM}} = 96$  eV
- $\tau_r \sim 200$   $\mu\text{s}$



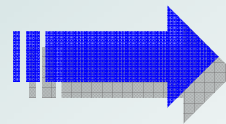
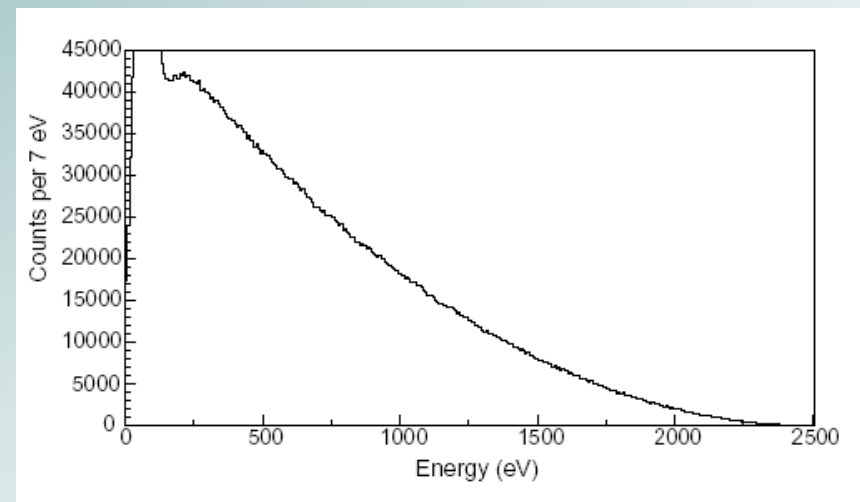
Total collected statistics:  
 $6 \times 10^6$   $\beta$  decays of  $^{187}\text{Re}$  above 420 eV

$$E_0 = 2470 \pm 1_{\text{stat}} \pm 4_{\text{sys}} \text{ eV}$$

$$\tau_{1/2} = 41.2 \pm 0.2_{\text{stat}} \pm 1.1_{\text{sys}} \text{ Gyr}$$

$$m_{\bar{\nu}_e}^2 = -462 \pm 579_{\text{stat}} \pm 679_{\text{sys}} (\text{eV})^2 / c^4$$

$$m_{\bar{\nu}_e} \leq 19.0 \text{ eV} / c^2 \text{ (90\% c.l.)}$$



Future improvements based on new  
 technology thermistors:  
**transition edge sensors (TES)**

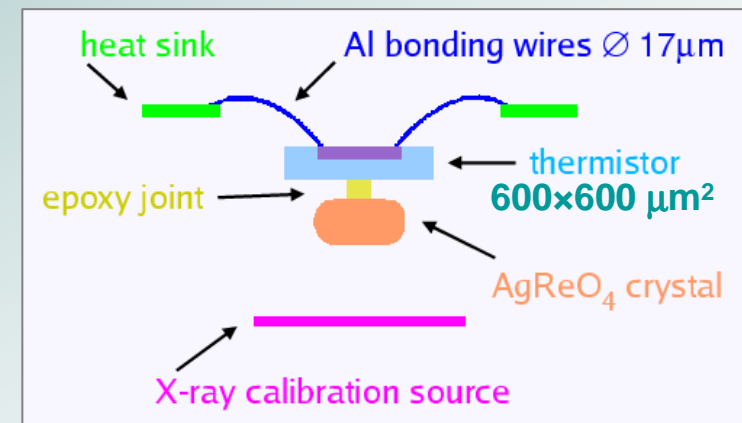
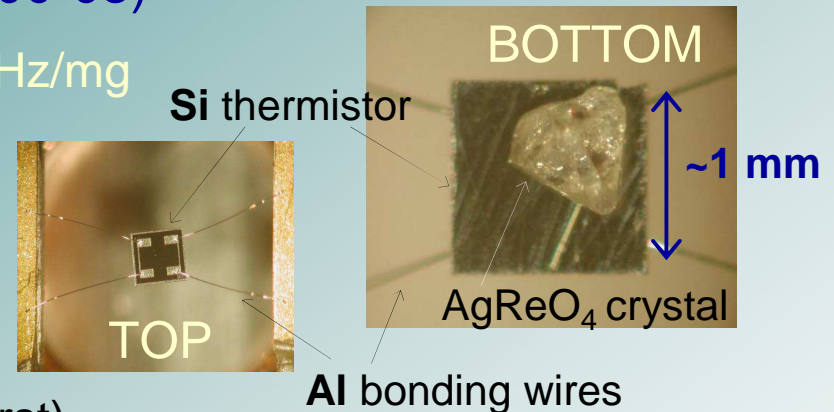
# Precursors experiments: MIBETA

(Milan-Como 2000-03)

- $\text{AgReO}_4$  single crystals  $\rightarrow$   $^{187}\text{Re}$  activity  $\cong 0.54$  Hz/mg
- Mass  $\cong 0.25$  mg  $\rightarrow A_\beta \cong 0.13$  Hz
  - ↳ to limit pile-up
- Array of 10 detectors
  - ↳ to increase statistics
- Phonon sensor: Si-implanted thermistors (ITC-irst)
  - ↳ high sensitivity
  - ↳ high reproducibility  $\Rightarrow$  arrays
  - ↳ possibility of  $\mu$ -machining

Technologies available for simultaneous fabrication of a large number, small dimension thermistors with fully integrated electrical connections

Reduced microphonism and problems of assembly



Useful for future expansion of arrays

# Results of MIBETA

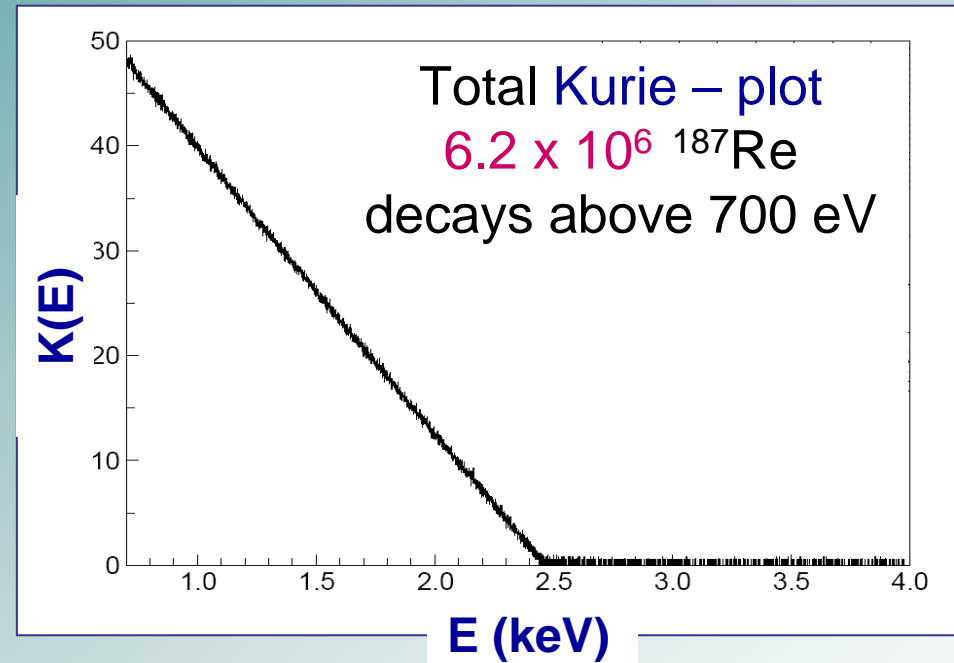
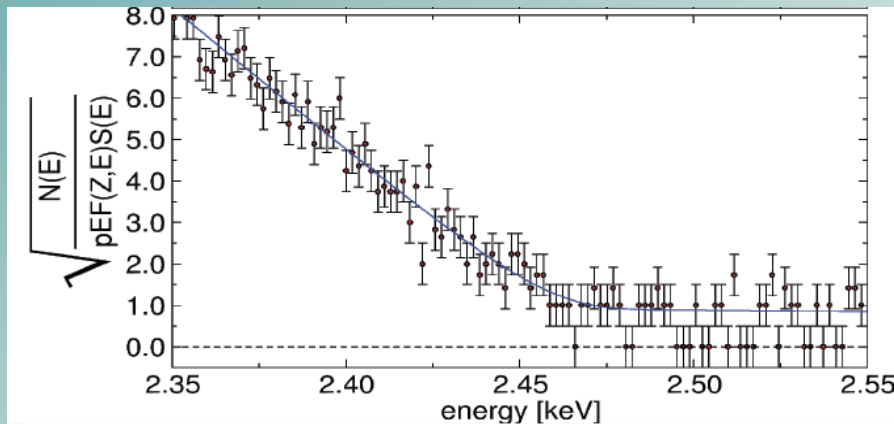
Total live time = 0.6 yr

Average  $\Delta E_{\text{FWHM}} = 28.5 \text{ eV}$

Average  $\tau_r = 490 \mu\text{s}$

$\beta$  spectrum fit

$$F = (f_{\text{th}} + f_{\text{pp}} + f_{\text{bck}}) \otimes f_{\text{det}}$$



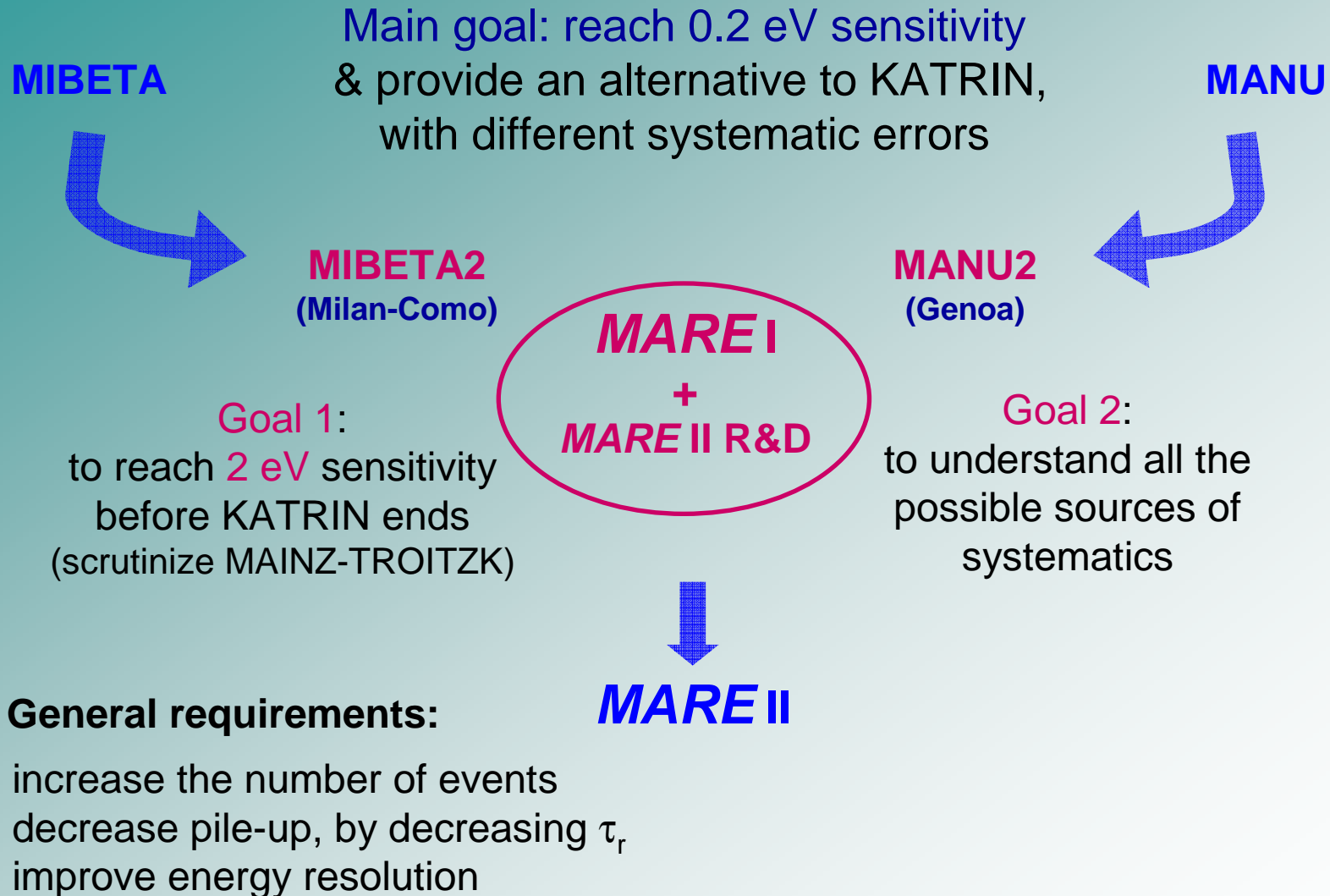
$$E_0 = 2465.3 \pm 0.5_{\text{stat}} \pm 1.6_{\text{sys}} \text{ eV}$$

$$\tau_{1/2} = 43.2 \pm 0.2_{\text{stat}} \pm 0.1_{\text{sys}} \text{ Gyr}$$

$$m_{\bar{\nu}_e}^2 = -112 \pm 207_{\text{stat}} \pm 90_{\text{sys}} (\text{eV})^2 / c^4$$

$$m_{\bar{\nu}_e} \leq 15.0 \text{ eV} / c^2 (90\% \text{ c.l.})$$

# The future



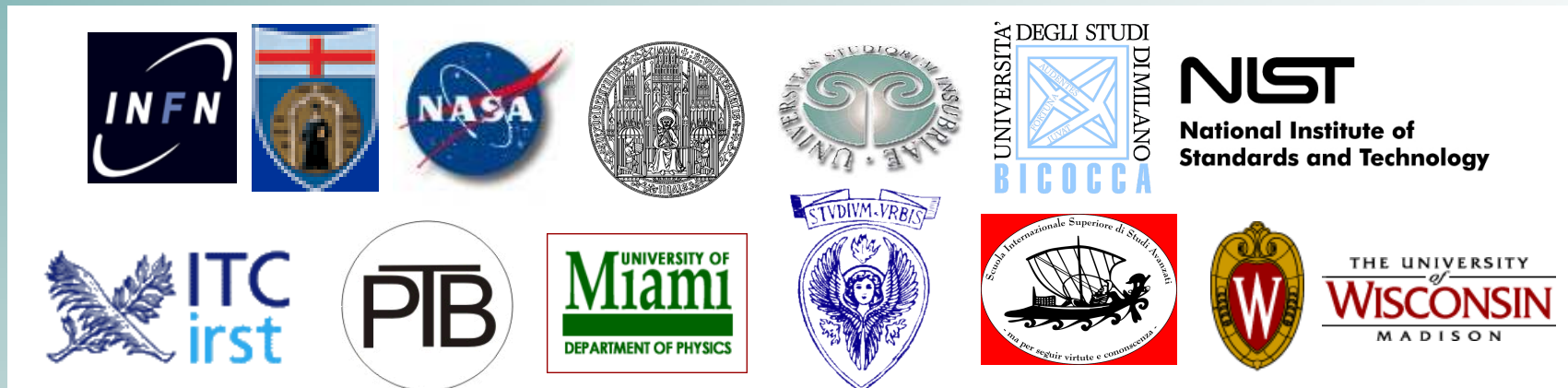


# The MARE collaboration

## MARE: Microcalorimeter Arrays for a Rhenium Experiment

Università di Genova, and INFN-Genova, Italy  
Goddard Space Flight Center, NASA, Maryland, USA  
Kirkhhof-Institute Physik, Universität Heidelberg, Germany  
Università dell'Insubria, and INFN-Milano-Bicocca, Italy  
Università di Milano-Bicocca, and INFN-Milano-Bicocca, Italy  
NIST, Boulder, Colorado, USA  
ITC-irst, Trento, and INFN-Padova, Italy  
PTB, Berlin, Germany  
University of Miami, Florida, USA  
Università di Roma "La Sapienza", and INFN-Roma1, Italy  
SISSA, Trieste, Italy  
Wisconsin University, Madison, Wisconsin, USA

<http://mare.dfm.uninsubria.it/frontend/exec.php>



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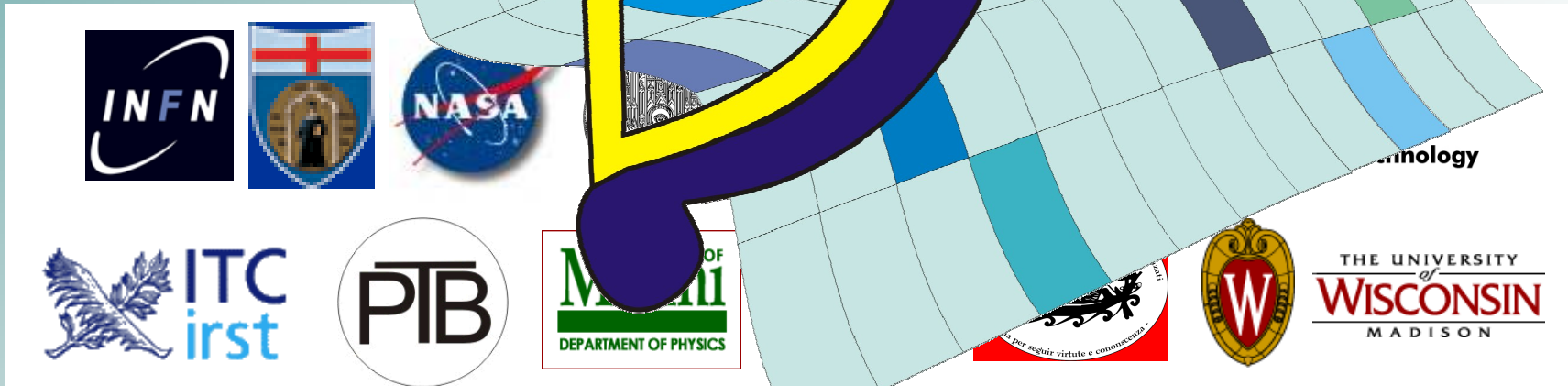


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Università di Pisa, INFN-Milano-Bicocca, Italy  
University of Colorado, USA  
Università di Padova, Italy  
University of Rome, Roma1, Italy

<http://mare.dfm.uninsubria.it/fronte>



Minneapolis, October 2008

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# MARE I

## 1. Reach 2 eV sensitivity:

- Present technology detectors
- Single channel optimization
- Scaling up to hundreds devices

$$\Delta E = 10 \text{ eV}, \tau_r = 150 \mu\text{s}$$

$$A_\beta = 0.3 \text{ Hz}, f_{pp} = 3 \times 10^{-5}$$

~ 300 detectors array

Total statistics ~  $10^{10}$  events

## GOALS

## 2. Improve understanding on systematics

- Theoretical spectral shape of decay
- Detector response function
- Unidentified pile-up
- ...

## 3. R&D for MARE II

## MONTECARLO simulations MARE I

Montecarlo input parameters			90% CL sensitivity	Possible experimental configurations			
$N_{ev}$ [ $\times 10^9$ ]	$f_{pile-up}$ [ $\times 10^{-5}$ ]	$\Delta E$ [eV]	$m_\nu$ [eV]	$N_{det}$	$t_M$ [y]	$\langle A_\beta \rangle$ [dec/s]	$\langle \Delta t \rangle$ [ $\mu\text{s}$ ]
1.4	2.0	10	3.5	100	2	0.20	100
3.2	2.5	10	3.0	200	2	0.25	100
4.7	2.5	10	2.5	200	3	0.25	100

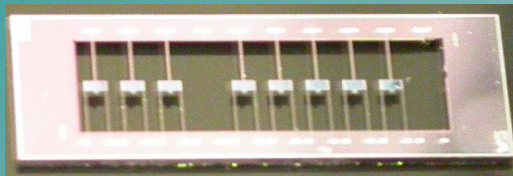
# MARE I: thermistors

$A_\beta \sim 0.3 \text{ Hz} \rightarrow$  doubling absorber mass while preserving  $\Delta E$  and  $\tau_r$

**Detectors require some improvements!**

## MIBETA2: available technologies

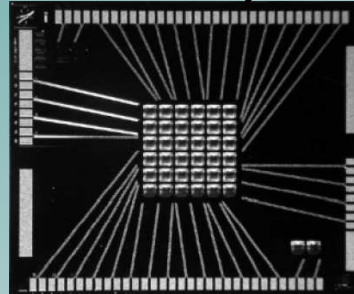
Arrays of 10 elements



Sensor area  $800 \times 800 \mu\text{m}^2$

**ITC-irst micromachined array.**  
Si-implanted produced by improving the technology developed for MIBETA.  
**Status:** ongoing production & tests; assembly problems

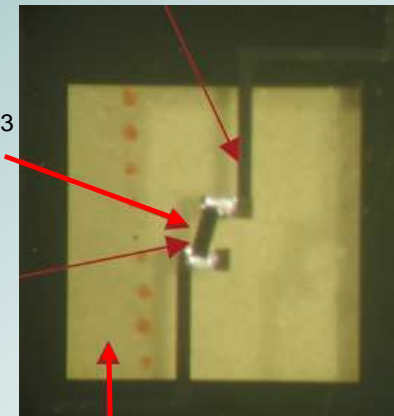
6x6 Si array



**NASA/GSFC 6x6 silicon array.**  
**Status:** encouraging first results. Coupling and electronics are being optimized.

NTD  
 $300 \times 200 \times 25 \mu\text{m}^3$

Nb electrical contact



SiN thermal link

**LBL+Bonn NTD Ge array.**  
**Status:** excess noise observed; reproducibility to be demonstrated.

## MANU2:

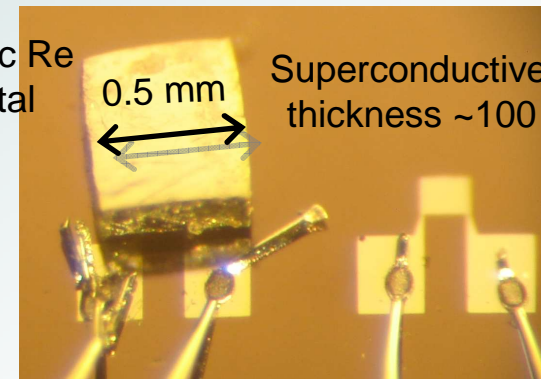
### Transition Edge Sensor (TES)

IrAu films instead of NTD thermistors  
» faster risetime and better S/N

Metallic Re crystal

0.5 mm

Superconductive film thickness  $\sim 100 \text{ nm}$





# MIBETA 2: previous tests

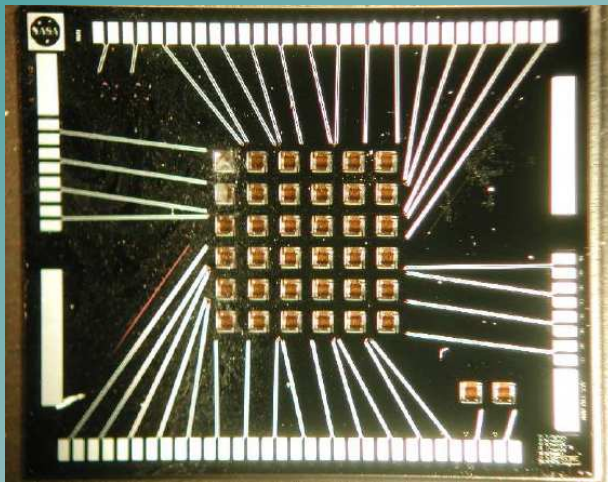
Array of **288 elements** achieved through a **gradual approach**

Single channel:

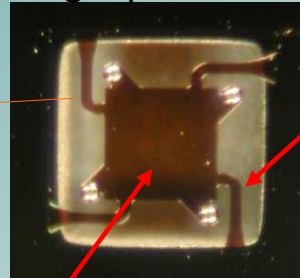
- $\text{AgReO}_4$  + semiconductor thermistor
- single crystal mass  $\sim$  **0.45 mg**  $\rightarrow$   $A_\beta \sim 0.3$  Hz

Basic structure:

NASA/GSFC XRS2: 6x6 Si-implanted array

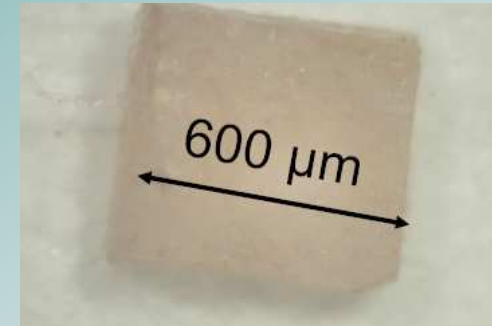


single pixel



Si-thermistor  
300x300x1.5  $\mu\text{m}^3$

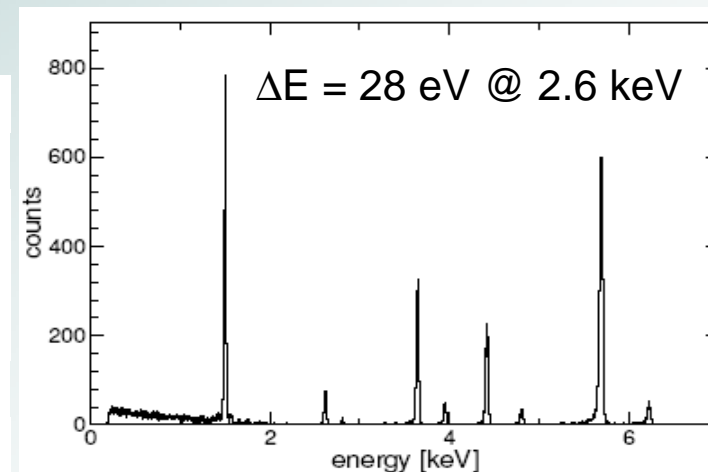
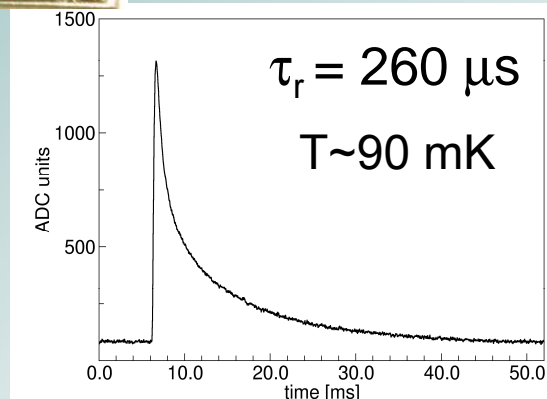
Si thermal  
links 3 $\mu\text{m}$



Already tested at  $\neq$  temperatures  
with glued crystals (in Milan)

Single channel best performances:

Further improvements  
attainable thanks to  
the new developed  
cold electronics  
(JFETs)



# MIBETA 2: schedule & sensitivity

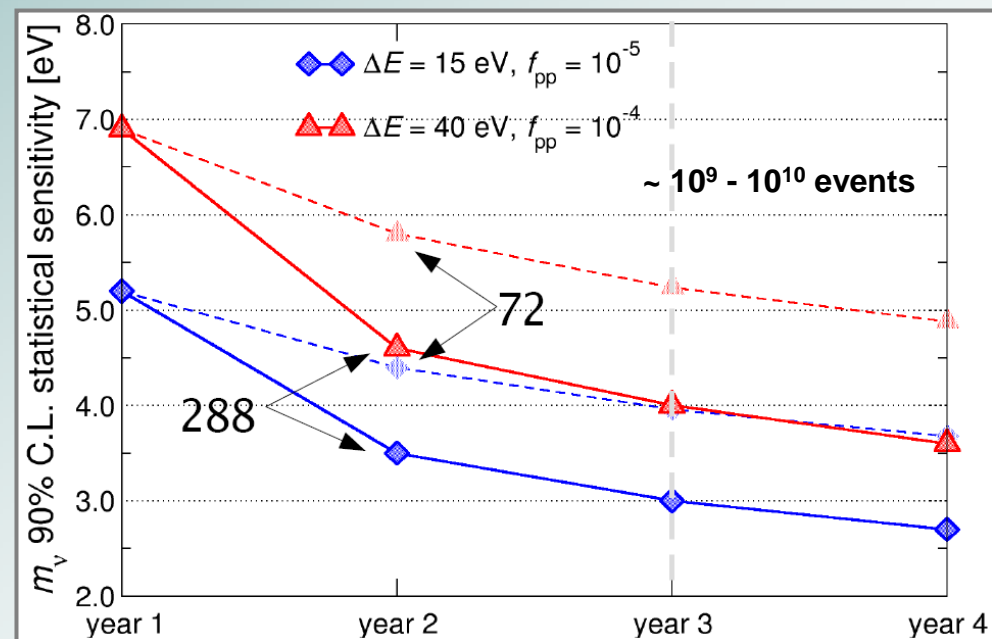
Gradual deployment of the whole 288 elements array:

year	1	2	3	4
new detectors	72	216	0	0
total detectors	72	288	288	288
statistics [det*y]	72	360	648	936
statistics [events]	$6.10 \times 10^8$	$3.05 \times 10^9$	$5.49 \times 10^9$	$7.94 \times 10^9$

- Array design based on XRS2 array
- Possible alternatives based on different (previously mentioned) thermistors
- Single channel activity:  $\sim 0.27$  Hz

Two approaches to evaluate the sensitivity:

- ❖ *conservative*:  
 $\Delta E = 40$  eV &  $\tau_r \sim 400$   $\mu$ s
- ❖ *improved*:  
 $\Delta E = 15$  eV &  $\tau_r \sim 50$   $\mu$ s

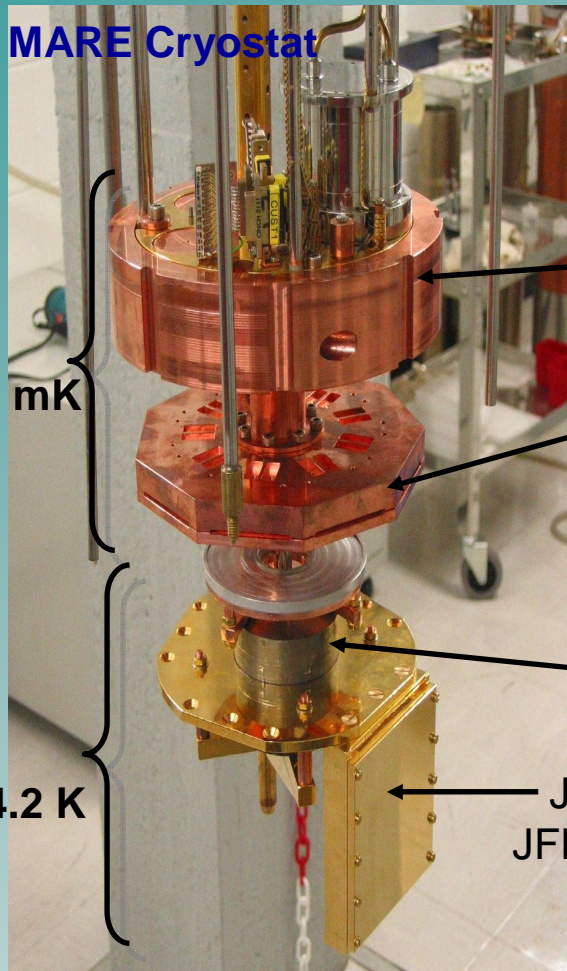




# MIBETA 2: status

- Cryogenic set-up (for 300 detectors) assembled in Milan
- 72 AgReO<sub>4</sub> crystals and 2 XRS2 arrays ready to be assembled
- Tests for better crystal-thermistor coupling already performed

MARE Cryostat



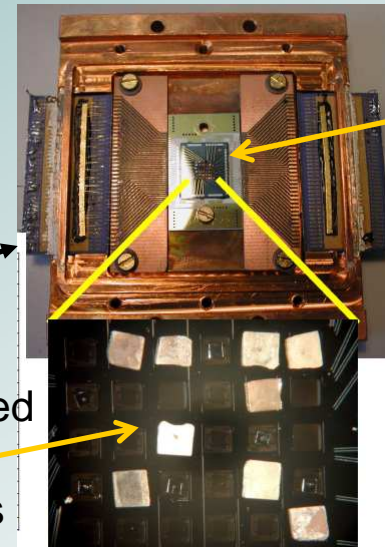
- Cold electronics is being assembled
- DAQ almost ready

Cu shield against environmental radioactivity

Array holder

Calibration X-ray source

JFET boxes (inside: JFETs heated to 150 K)



XRS2 array

Crystals glued on top of thermistors

Data taking with first **72** channels expected to start at the **end 2008 - beginning 2009**

# MARE II

## GOAL

reach 0.2 eV sensitivity around 2015

### Requirements (from MonteCarlo simulations)

Total statistics ~  $10^{14}$  events

Activity/element ~ 1-10 Hz

$\tau_r \sim 1 - 10 \mu\text{s}$

$\Delta E_{\text{FWHM}} \sim 5 \text{ eV}$

~ 4 years R&D

#### Kick-off of MARE II subordinated to:

- safe reduction of known sources of systematics;
- verification that no new sources appear;
- complete understanding of the  $^{187}\text{Re}$  decay spectrum;
- demonstration that the estimated sensitivity can be maintained through the experiment segmentation & expansion

#### Substantial improvements are needed:

- sensors: TES or MMC or MKID
- electronics: multiplexed SQUID
- methods: modularity

→ scaling up to **thousands** devices!

Technologies **already under study**  
in several other experiments

The full **MARE phase I** dataset is required to draw a definitive conclusion.

# MARE II

**Thermometer:** a rise time of  $\sim 1 - 10 \mu\text{s}$  is required  
higher statistics with lower pile-up

## Multiplexed kinetic inductance detectors (MKIDS)

(Roma, ITC-irst, Cardiff)

Superconductive strip below  $T_c$  whose surface inductance  $L_s$  and impedance  $Z_s$  are changed by absorption of quasi particles; the signal is read as a phase variation when the strip is part of a resonant circuit

## Magnetic MicroCalorimeters (MMC)

(Heidelberg)

Paramagnetic material in a small magnetic field with temperature dependent magnetization

## Transition Edge Sensors (TES)

(MANU 2)

Already mentioned: Temperature sensitivity  $\sim 60$  times larger than for doped semiconductor thermistors; Metallic  $\rightarrow$  e-ph coupling time shorter than for doped semiconductors

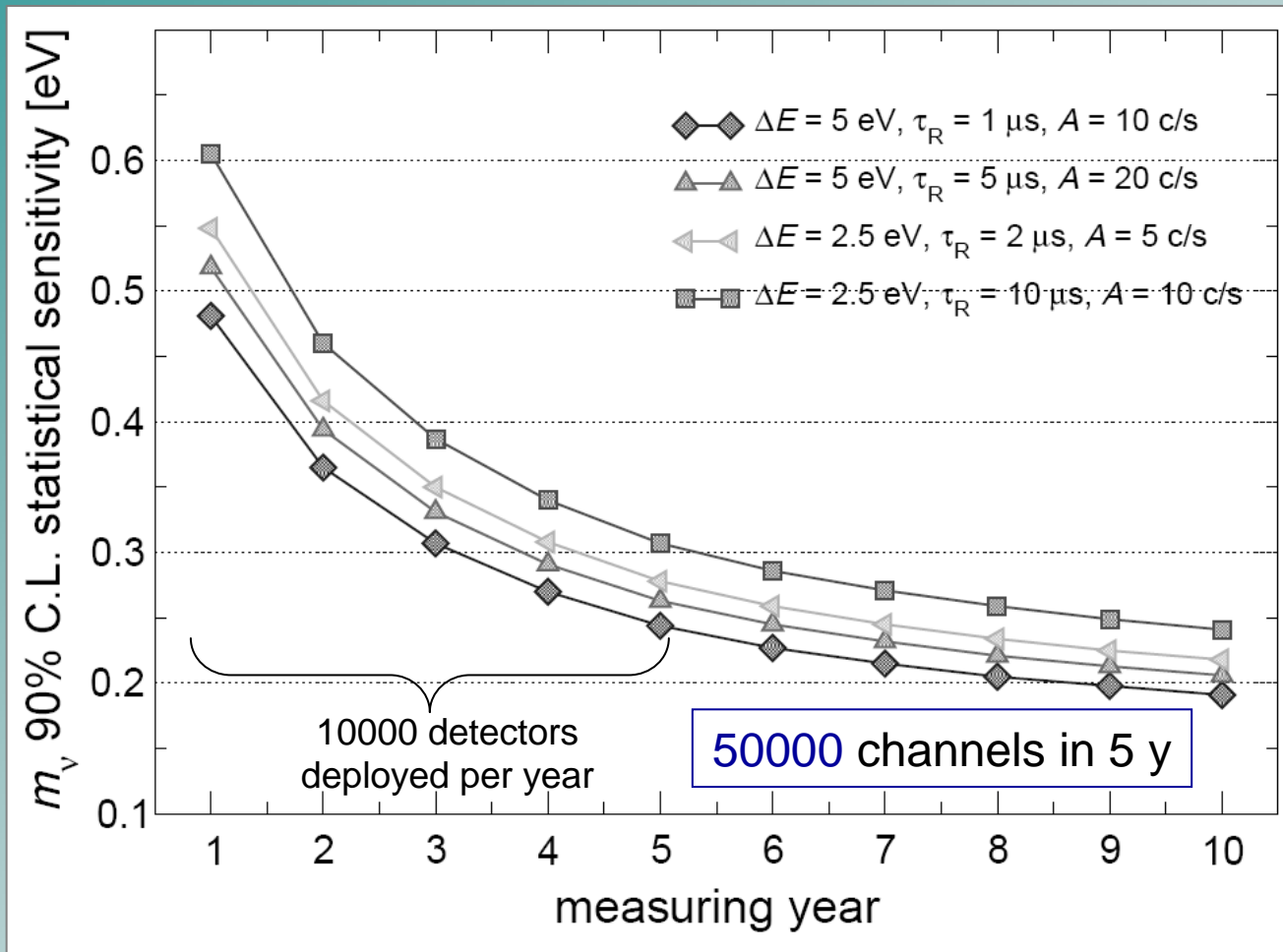
## Electronics: front-end multiplexed SQUID (NIST, PTB)

Very good noise performance allows construction of multiplexers that read out a number of sensors on a single channel; couples very well to TES, MMC, MKIDS

# MARE II

Approach: design a kind of modular 10000 pixel array kit which can be relatively easily installed in any available refrigerator

## Simulations MARE II



10000 detectors  
deployed per  
year



# Conclusions

- The **calorimetric technique** can reach sub-eV sensitivity on  $m_\nu$  being complementary to KATRIN
- The **MARE** experiment will be developed into 2 phases:
  - **MARE I**: important to understand all sources of systematics by implementing the specific know-how developed by the involved groups
  - **MARE II**: new technology thermistors & read-out are needed to achieve the experimental requirements
- Thanks to the **modularity** of the calorimetric approach a further expansion of the experiment will simply consist in the repeated replication of the first matrix (unlike spectrometers)

**Thanks!**

