

# Mini-CLEAN and DEAP/CLEAN Detectors

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Dark Side of the Universe  
University of Minnesota

# The Noble Liquid Revolution

Noble liquids are relatively inexpensive, easy to obtain, and dense.

Easily purified

- low reactivity
- impurities freeze out
- low surface binding
- purification easiest for lighter noble liquids

Ionization electrons may be drifted through the heavier noble liquids

Very high scintillation yields

- noble liquids do not absorb their own scintillation
- 30,000 to 40,000 photons/MeV
- modest quenching factors for nuclear recoils

Easy construction of large, homogeneous detectors

# Liquified Noble Gases: Basic Properties

Dense and homogeneous

Do not attach electrons, heavier noble gases give high electron mobility

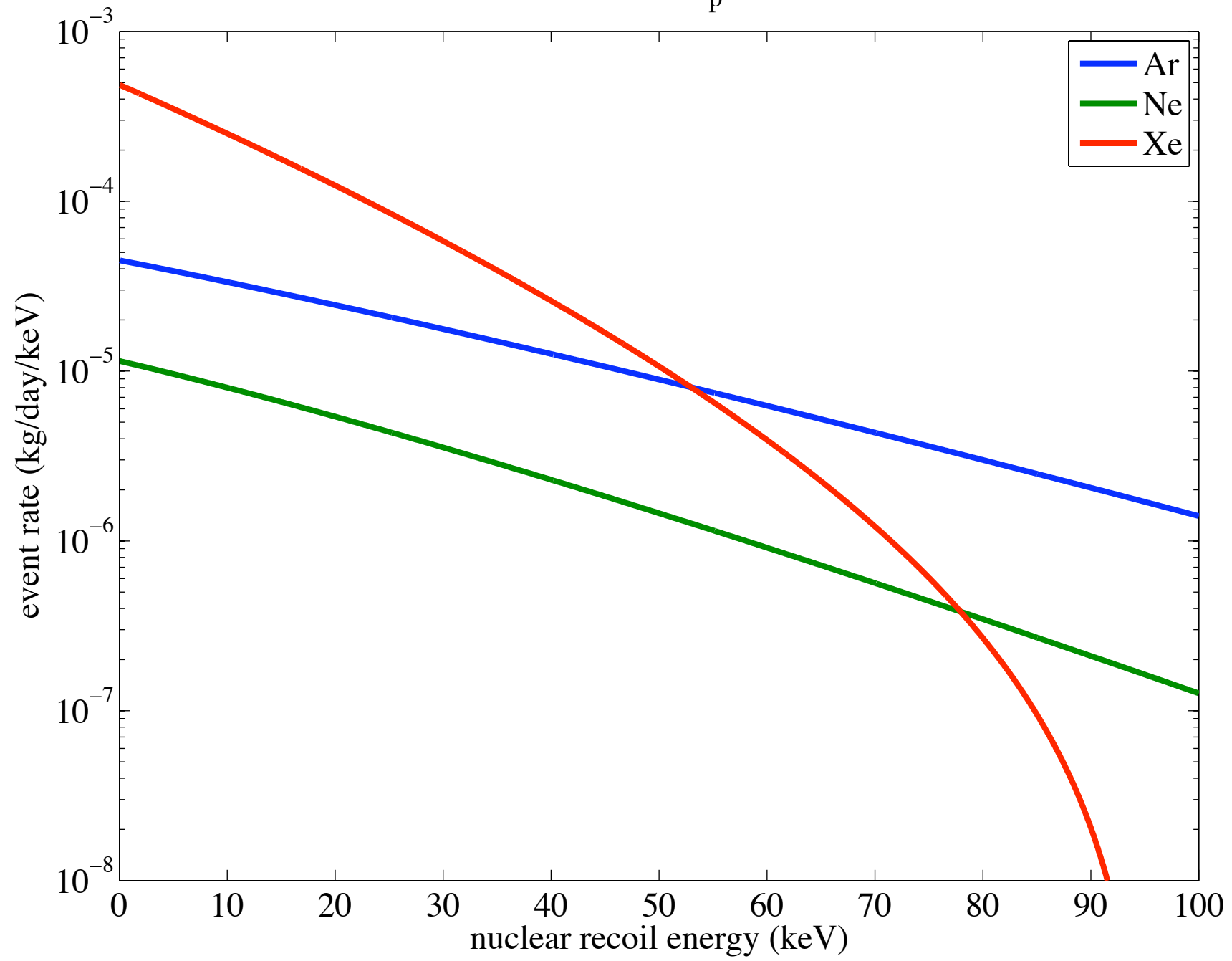
Easy to purify (especially lighter noble gases)

Inert, not flammable, very good dielectrics

Bright scintillators

	Liquid density (g/cc)	Boiling point at 1 bar (K)	Electron mobility (cm <sup>2</sup> /Vs)	Scintillation wavelength (nm)	Scintillation yield (photons/MeV)	Long-lived radioactive isotopes	Triplet molecule lifetime (μs)
LHe	0.145	4.2	low	80	19,000	none	13,000,000
LNe	1.2	27.1	low	78	30,000	none	15
LAr	1.4	87.3	400	125	40,000	<sup>39</sup> Ar, <sup>42</sup> Ar	1.6
LKr	2.4	120	1200	150	25,000	<sup>81</sup> Kr, <sup>85</sup> Kr	0.09
LXe	3.0	165	2200	175	42,000	<sup>136</sup> Xe	0.03

100 GeV WIMP  $\sigma_p = 10^{-44} \text{ cm}^2$



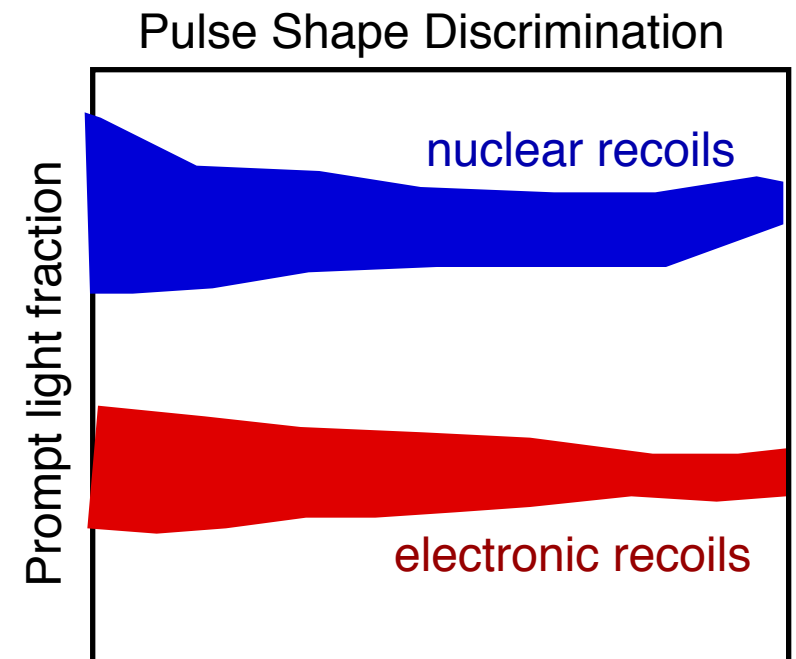
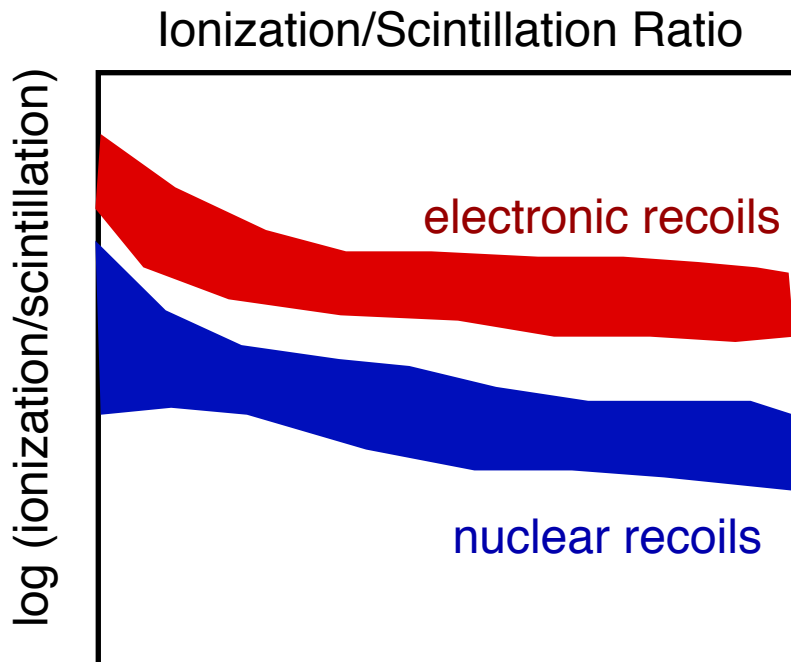
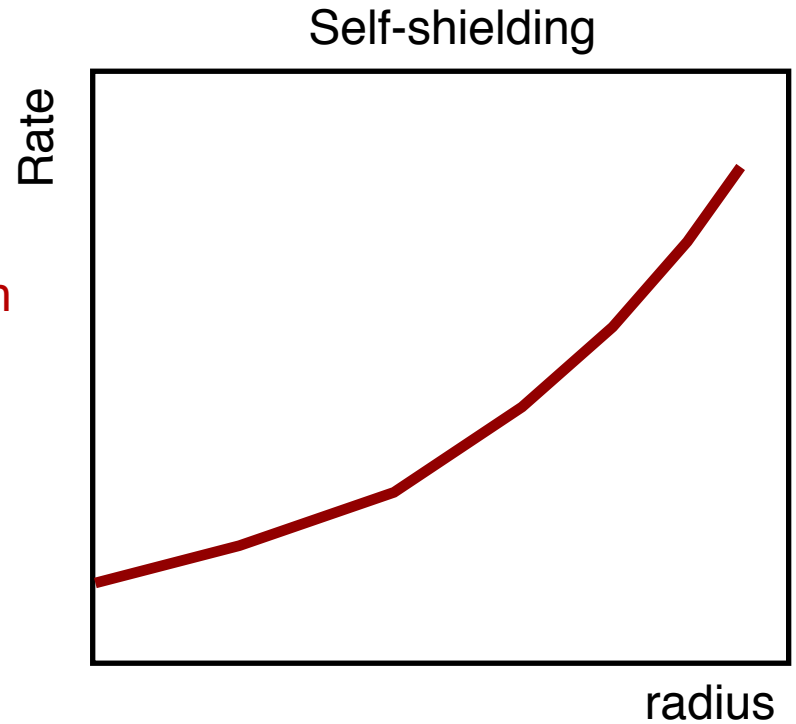
# Strategies for Electronic Recoil Background Reduction in Scintillation Experiments

Require  $< 1$  event in signal band during WIMP search

LXe: Self-shielding, Ionization/Scintillation ratio best

LAr: Pulse shape, Ionization/Scintillation ratio best

LNe: Pulse shape, Self-shielding best



# The Mini-CLEAN Approach

Scaleable technology based on detection of scintillation in liquified noble gases. No E field. Ultraviolet scintillation light is converted to visible light with a wavelength-shifting film.

Liquid neon and liquid argon are bright scintillators (30,000 - 40,000 photons/MeV).

Do not absorb their own scintillation.

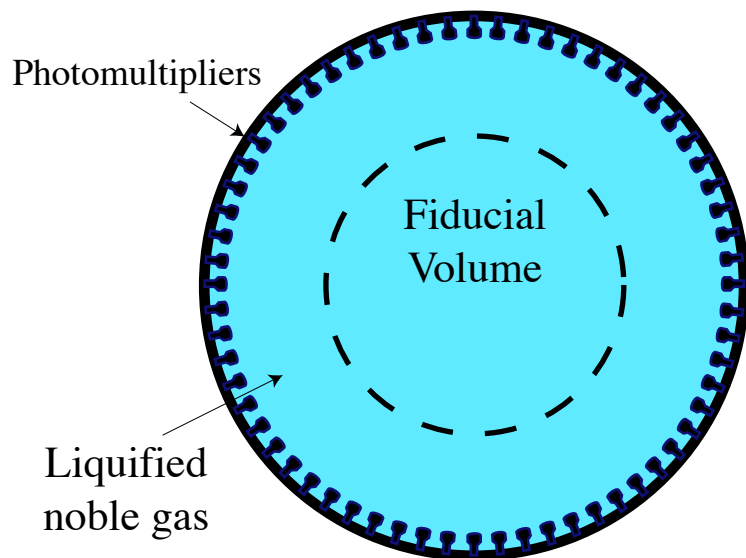
Are inexpensive (Ar: \$2k/ton, Ne: \$60k/ton).

Are easily purified underground.

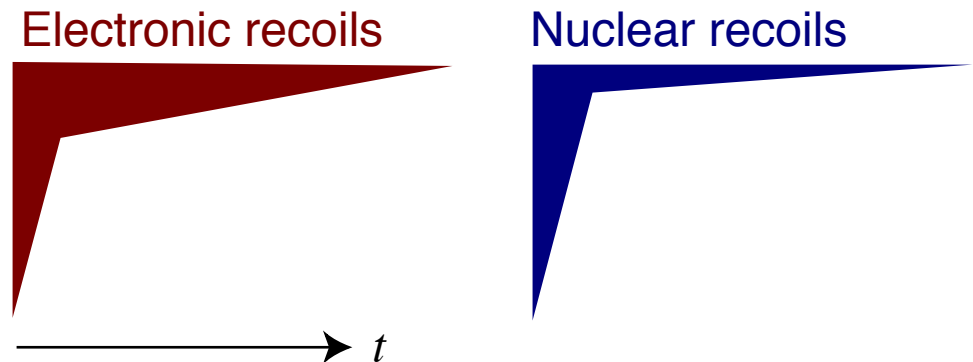
Exhibit effective pulse shape discrimination.

Exchange of targets allows better characterization of radioactive backgrounds

## Self-shielding



## Pulse-shape discrimination



Fast component:  $< 10$  ns

Slow component: 1.6  $\mu$ s (LAr), 15  $\mu$ s (LNe)

Discriminate based on fraction of light in first 100 ns (Fprompt)

D. N. McKinsey and J. M. Doyle, J. Low Temp. Phys. 118, 153 (2000).

D. N. McKinsey and K. J. Coakley, Astropart. Phys. 22, 355 (2005).

M. Boulay, J. Lidgard, and A. Hime, nucl-ex/0410025

M. Boulay and A. Hime, Astropart. Phys. 25, 179 (2006).

# The CLEAN/DEAP family

A series of detectors based entirely on scintillation in LNe, LAr

No electric field!

This allows very efficient light collection with 4 $\pi$  photodetector coverage

✓ Pico-CLEAN: 200 g of LNe or LAr. Completed LNe scintillation R&D.

✓ Micro-CLEAN: 4 kg of LNe or LAr. LAr and LNe R&D underway.

DEAP-I: 7 kg LAr. DM search at SNOLAB. To go underground in 2007.

Mini-CLEAN: 100 fiducial kg of LNe or LAr. R&D, DM search. Begin 2007?

CLEAN/DEAP: 1 fiducial ton LAr/LNe. DM search. 2009 - ?

CLEAN: 10 fiducial tons LNe, maybe LAr. pp neutrinos, DM search. 2011 ?

# CLEAN/DEAP Collaboration

*Yale University:* W. H. Lippincott, D. McKinsey, J. Nikkel

*Boston University:* D. Gastler, E. Kearns

*Carleton University:* K. Graham, K. McFarlane

*Los Alamos National Laboratory:* A. Hime, F. Lopez, J. Oertel, L. Rodriguez, K. Rielage,  
L. Stonehill, J. Wouters

*National Institute for Standards and Technology:* K. J. Coakley

*Queen's University:* M. Boulay, A. Hallin, J. Lidgard, R. Matthew, A. McDonald, P. Skensved

*SNOLAB:* F. Duncan, C. J. Jillings, I. Lawson

*University of North Carolina:* R. Henning

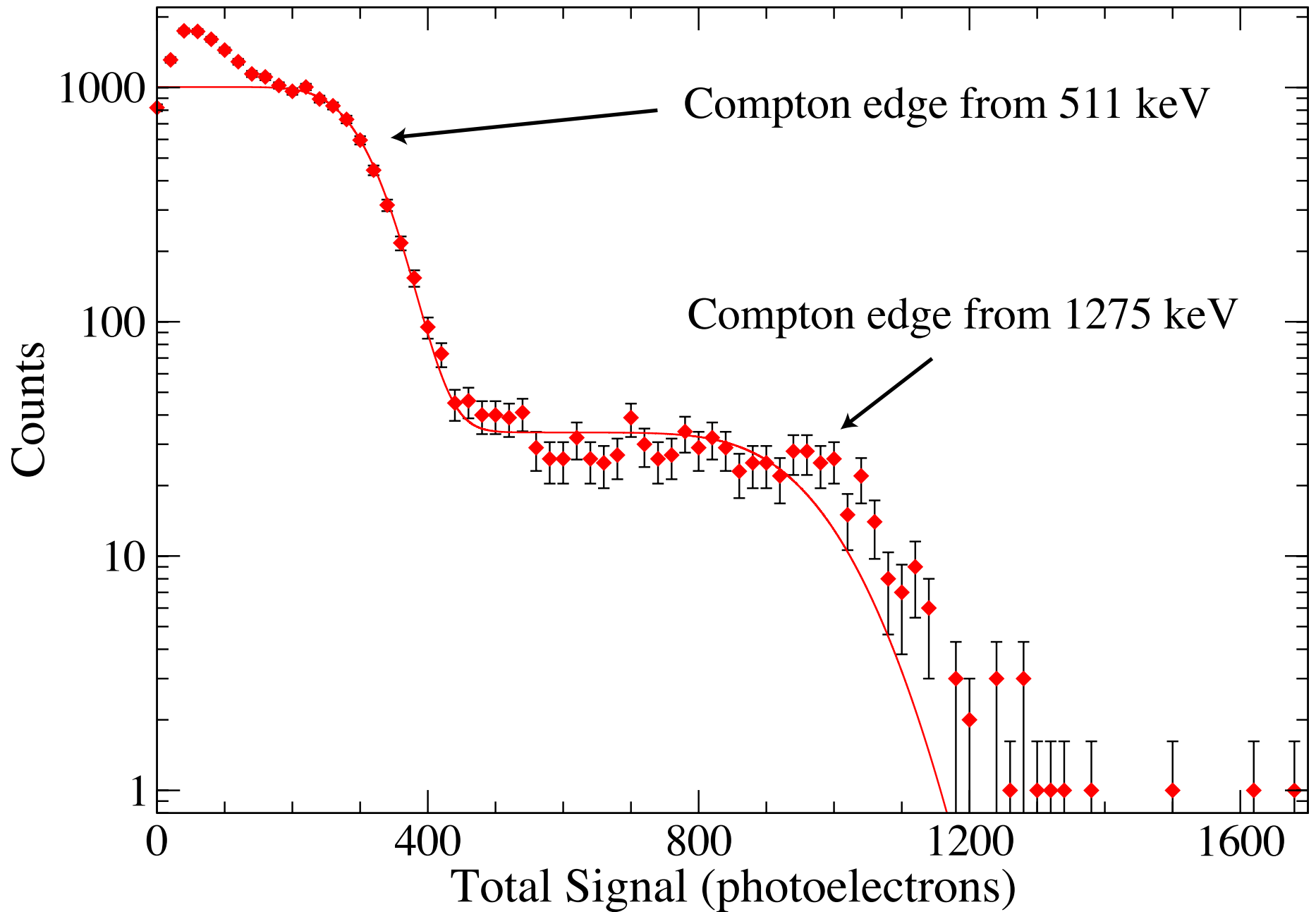
*University of South Dakota:* D. Mei

*University of Texas:* R. Hegde, J. Klein, S. Seibert

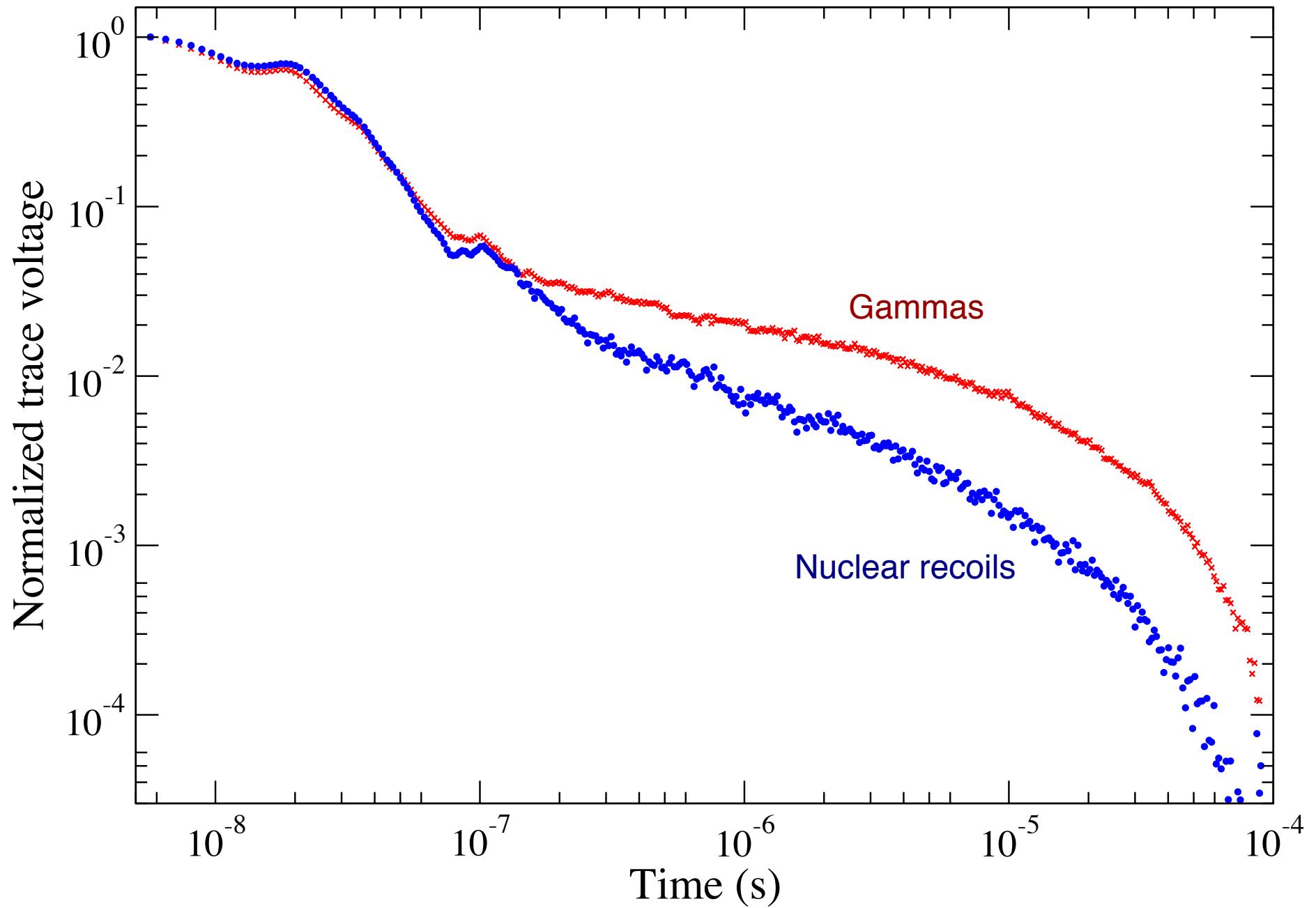


Liquid Neon calibration with Na-22 gamma rays gives  $0.9 \text{ pe/keV}$

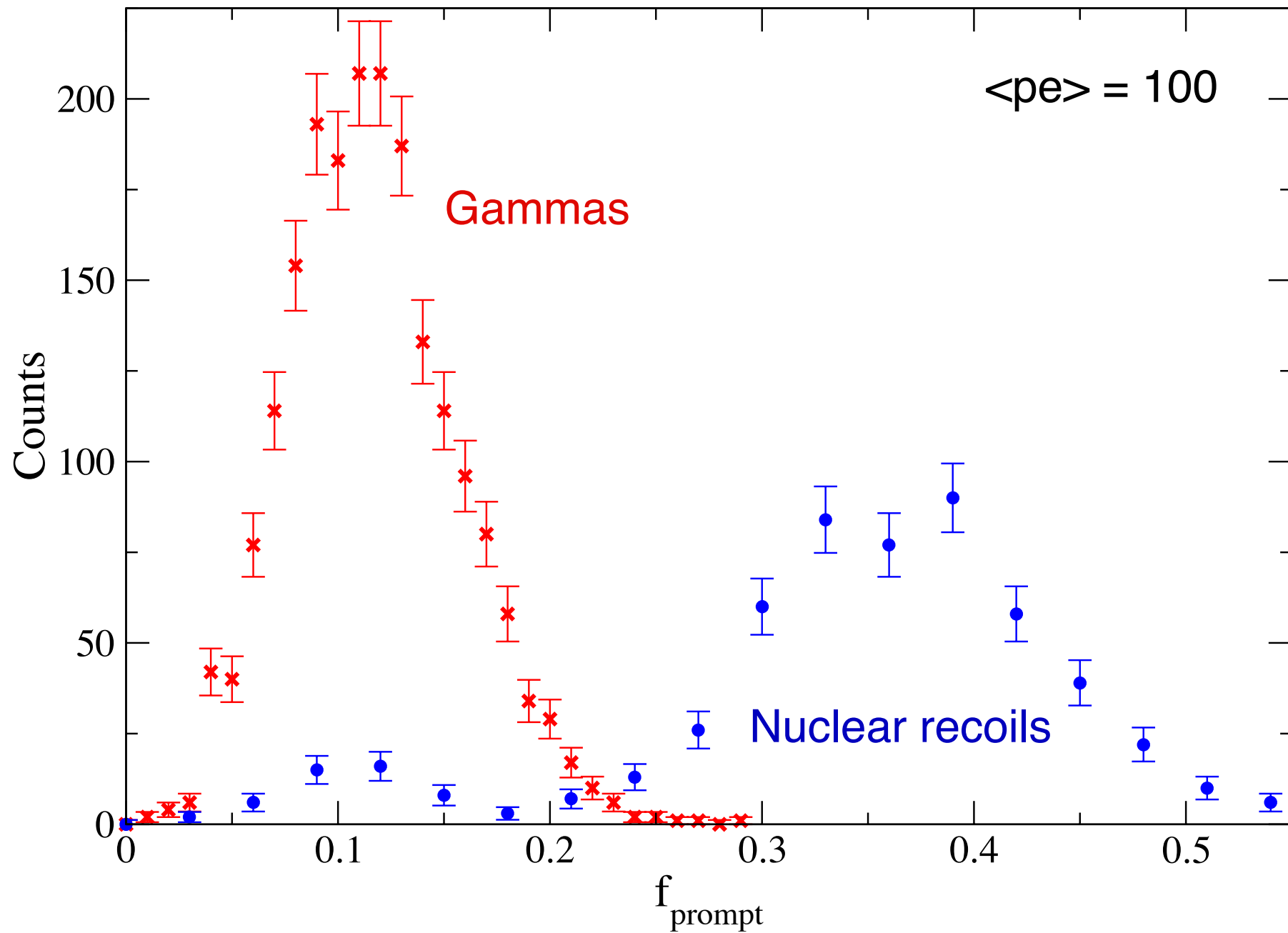
→  $30,000 \text{ photons/MeV}$



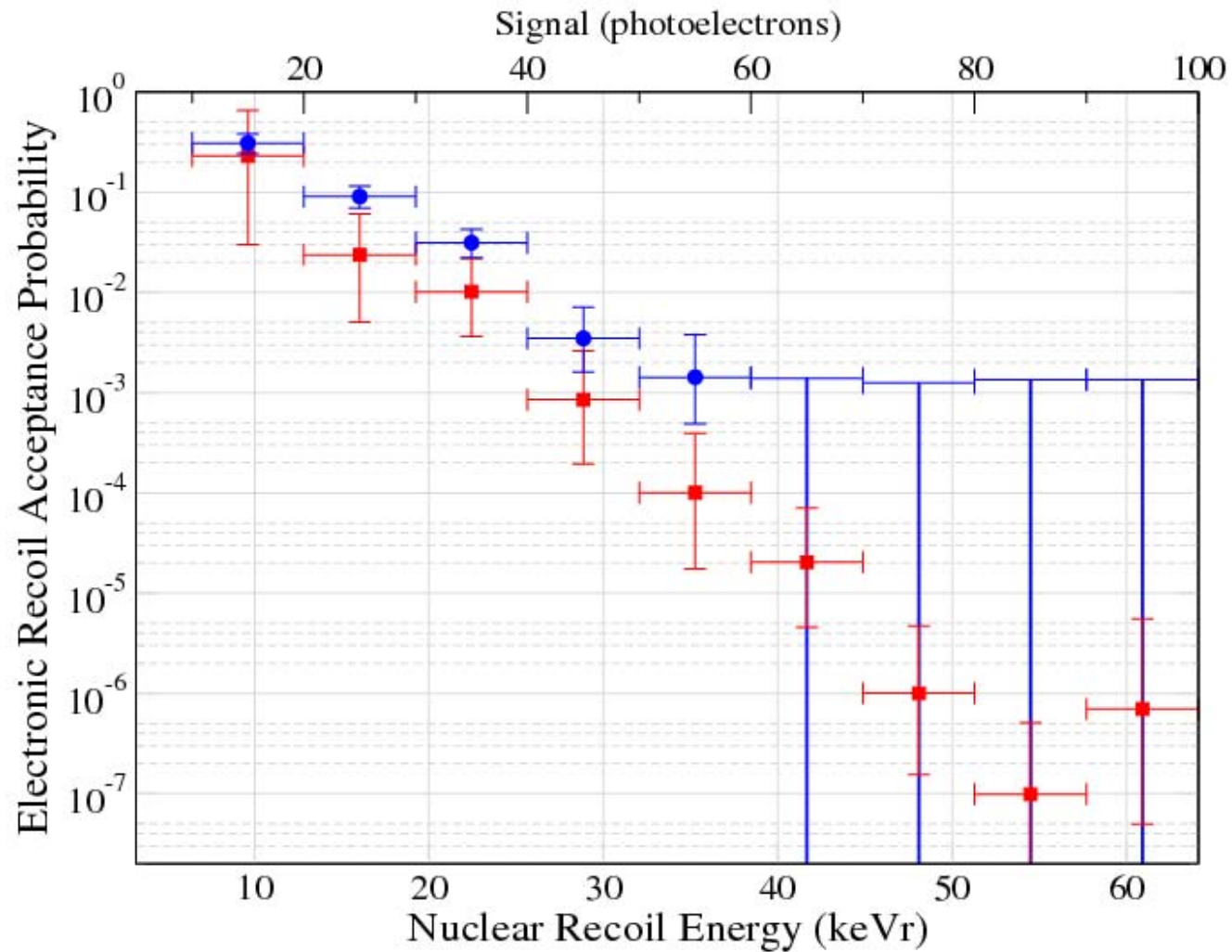
# Scintillation Time Dependence in LNe



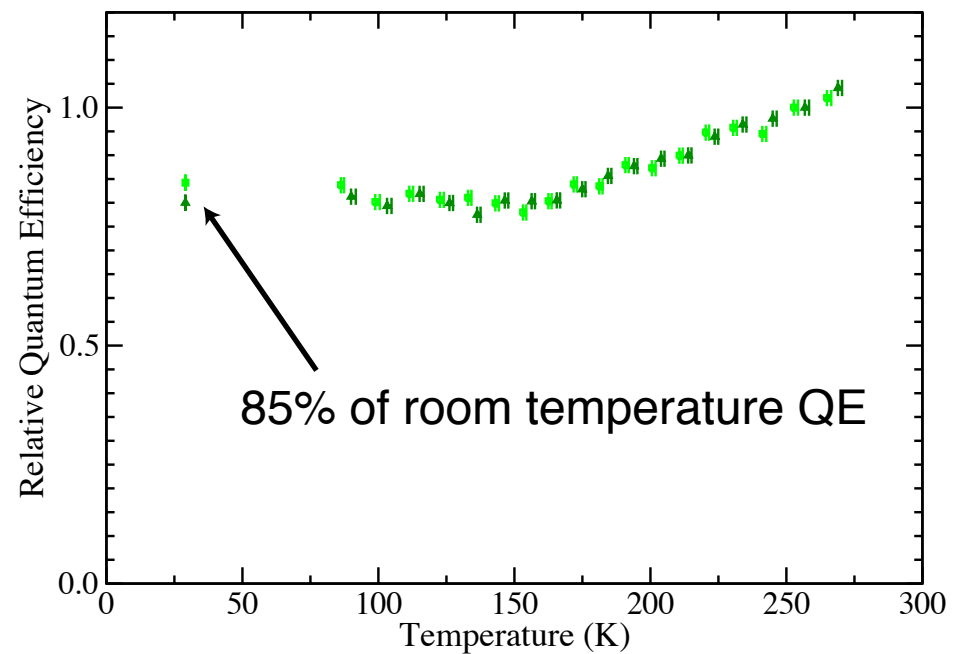
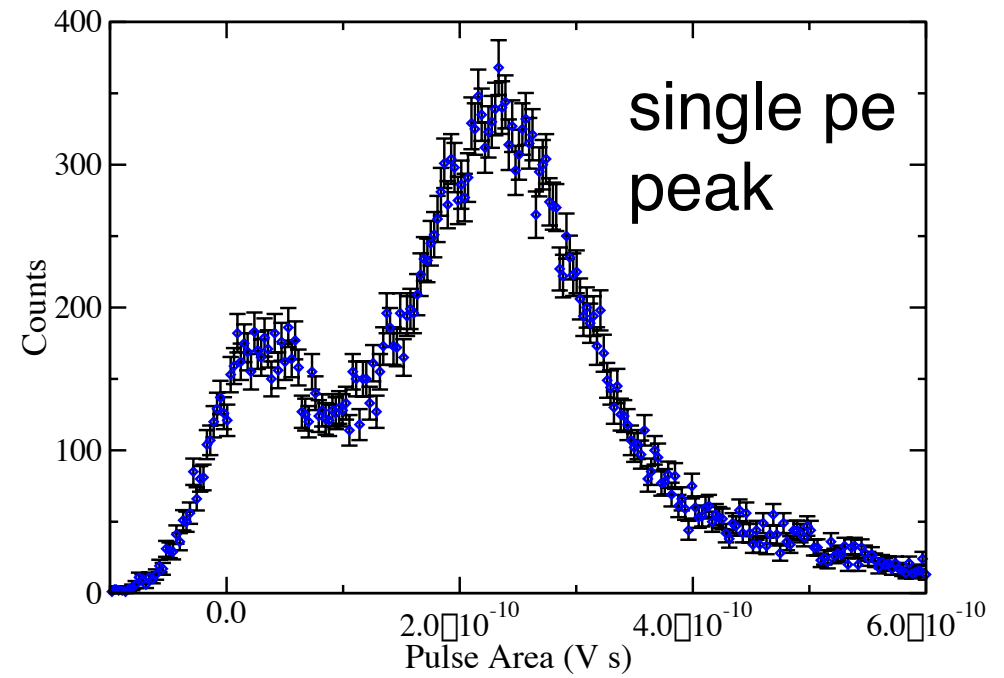
# Pulse shape discrimination in LNe



# Neon PSD from pico-CLEAN



# PMT testing at 27 K



# Micro-CLEAN



Currently in operation at Yale  
Two 20-cm PMTs  
4 kg active mass of LAr or LNe  
PMTs are immersed in the liquid

Signal:  
4.2 pe/keV for electronic recoils in 2006.

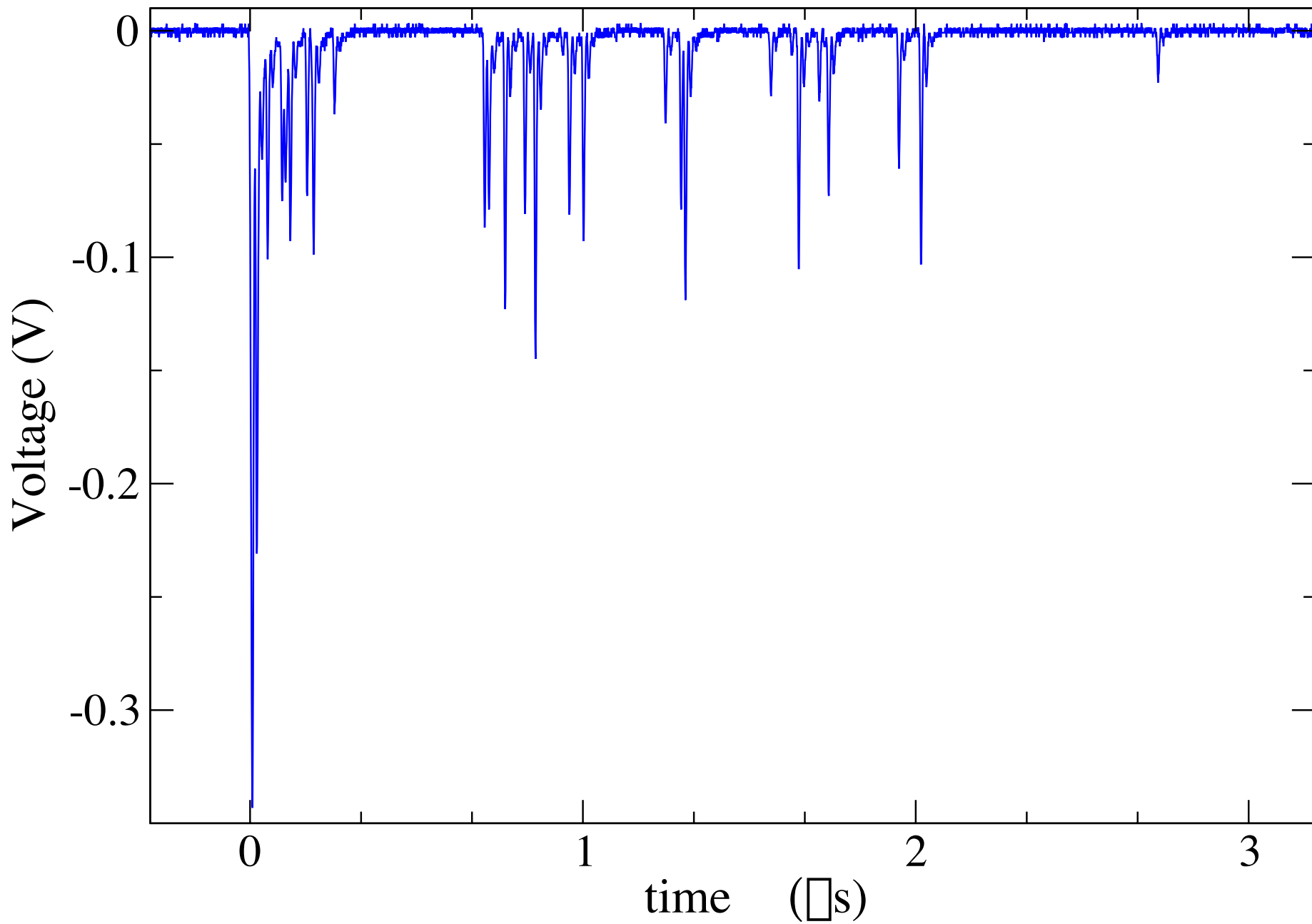
Jan-Feb 2007: improved WLS coating  
added continuous purification

-----> 4.9 pe/keV right now

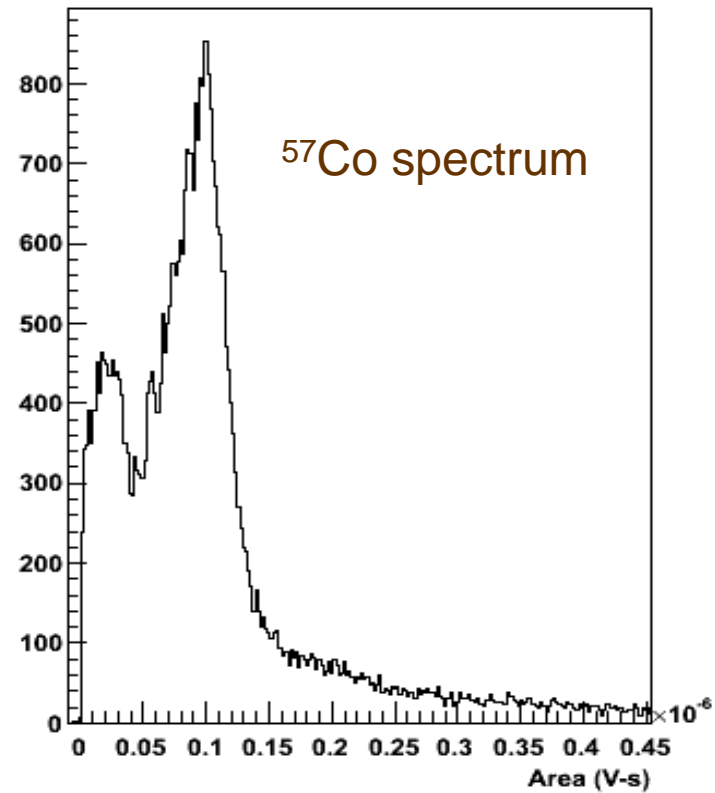
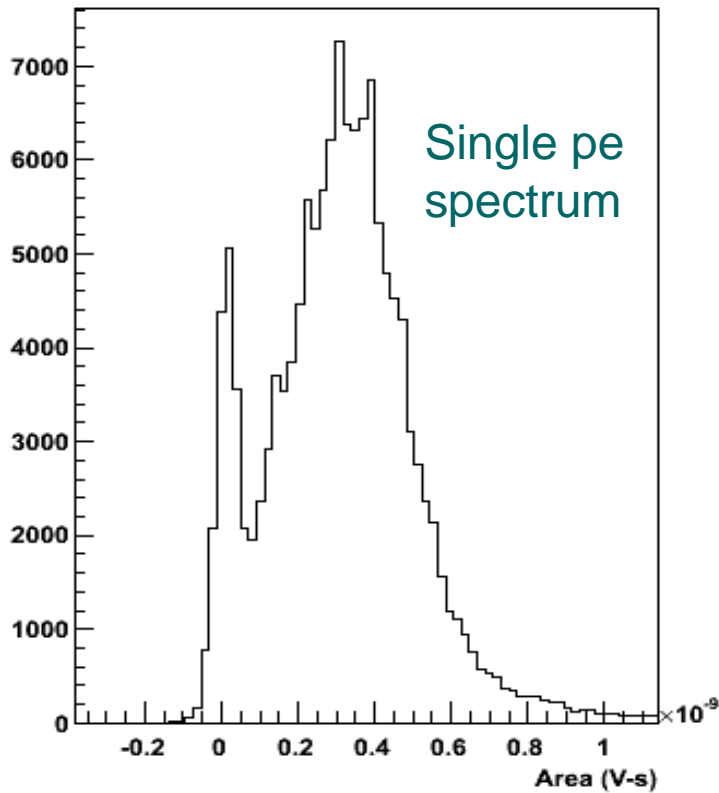
Goals:

- 1) Measure light yield for nuclear recoils
- 2) Test PMTs in LAr, LNe
- 3) Test digitization electronics  
(from Boston U)
- 4) Test discrimination down to 10 keVr

# Sample scintillation pulse (gamma Compton scatter in LAr)



# Light Yield

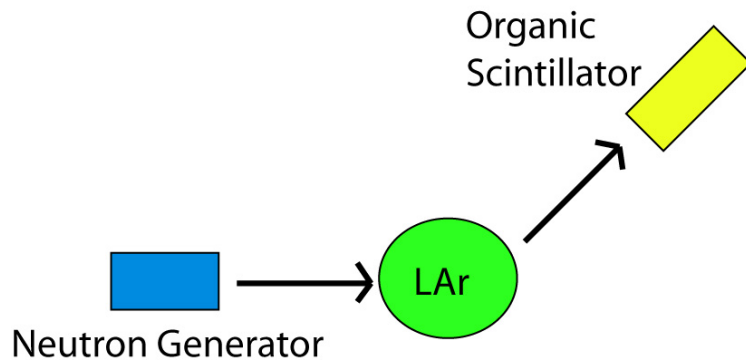


- Single photoelectron spectrum and cobalt peak for one of the PMTs
- Measure the light yield to be 4.9 pe/keVee

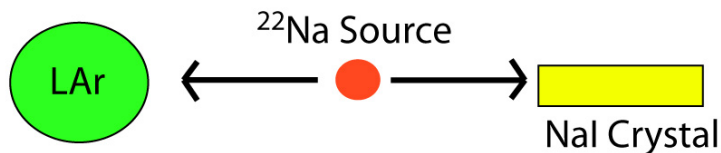


# Data Tagging and Quality

## Nuclear Recoil Data

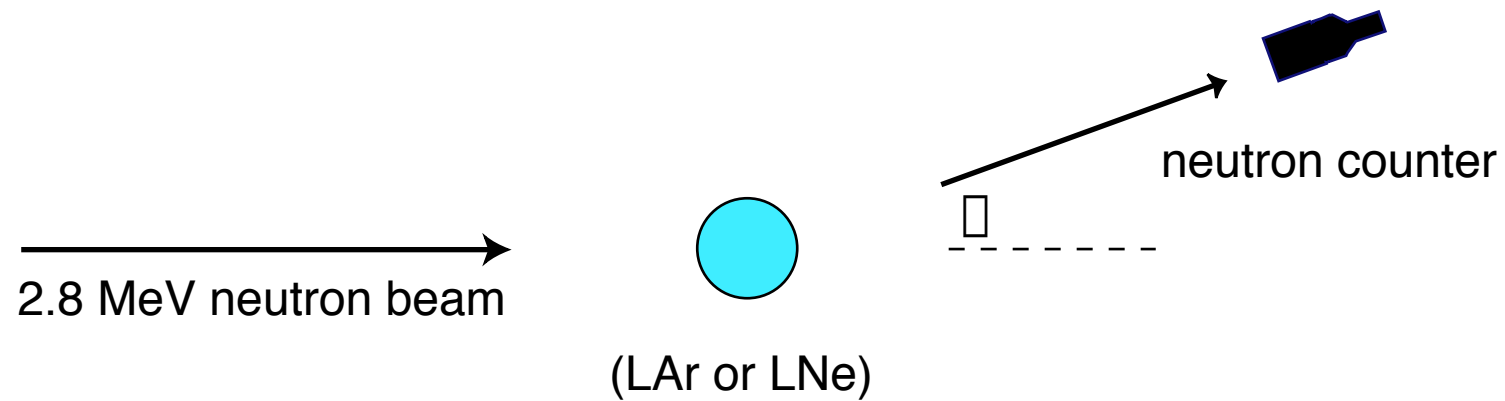


## Electronic Recoil Data



- Require coincidence with external detector
- Time of Flight cut
- PSD cut in organic scintillator
- Energy cut in NaI crystal
- Asymmetry cut in the PMTs

# Calibration of nuclear recoil-induced scintillation in LAr and LNe

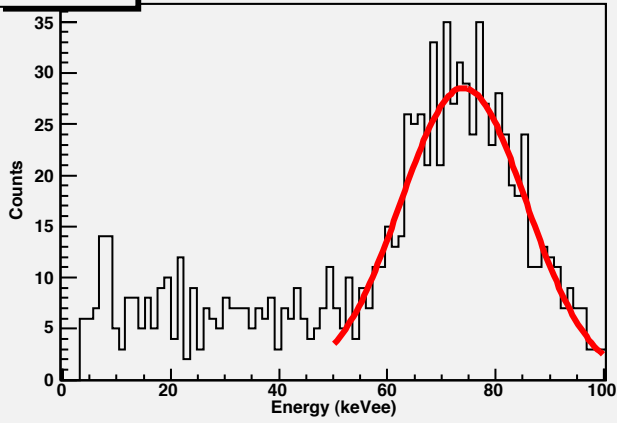


Require delayed coincidence between cryogenic scintillator and neutron counter

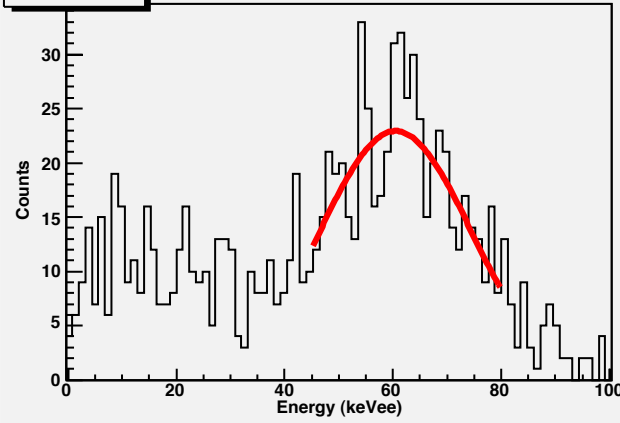
$$E_A = E_n \frac{2M_n M_A}{(M_n + M_A)^2} (1 - \cos \theta)$$

# Neutron Peaks at a Variety of Scattering Angles in LAr

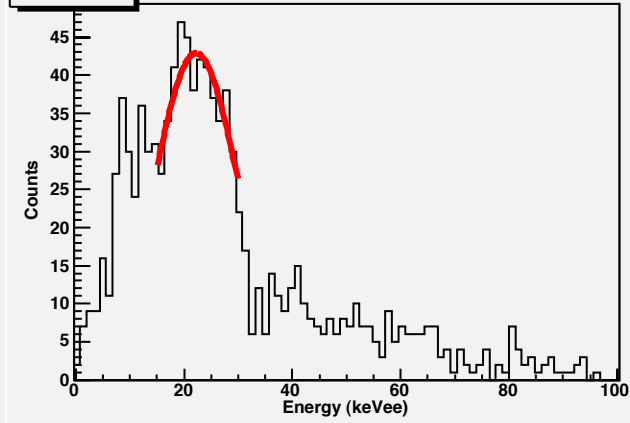
238.6 keV



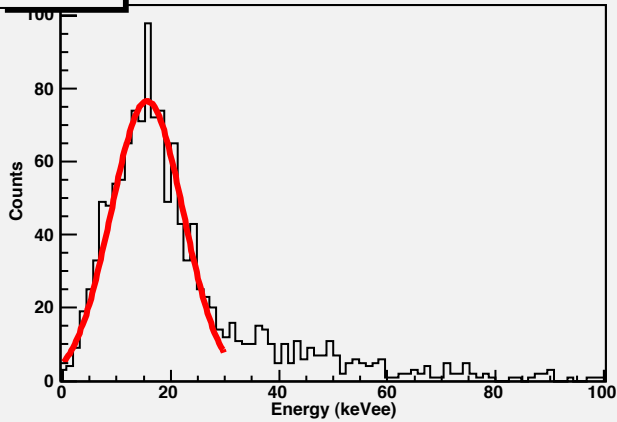
210.1 keV



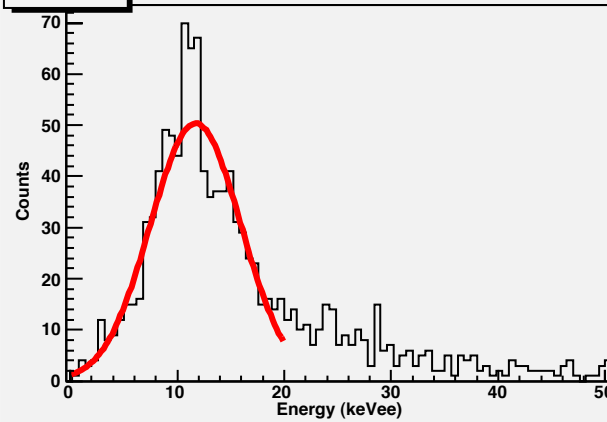
86.2 keV



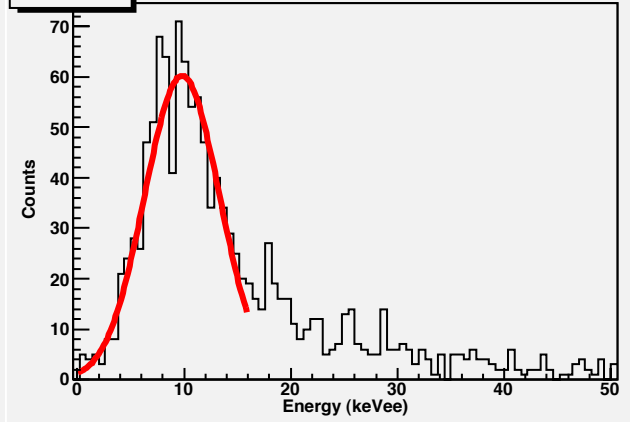
66.9 keV



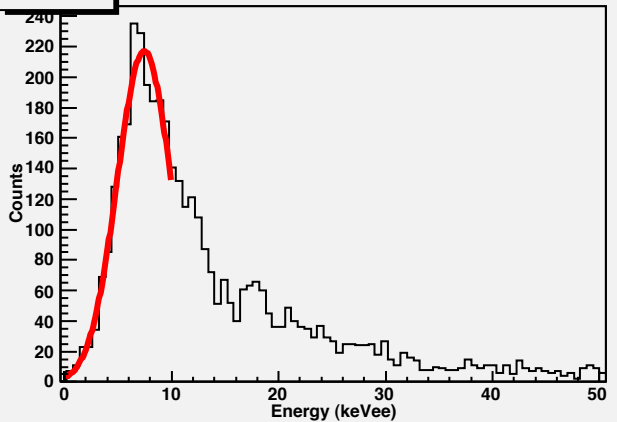
44.5 keV



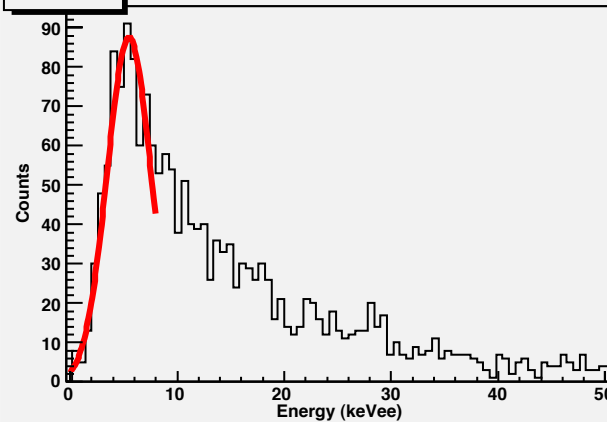
32.6 keV



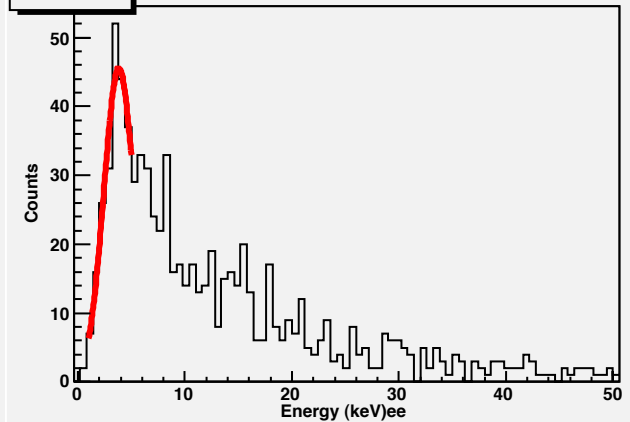
27.7 keV



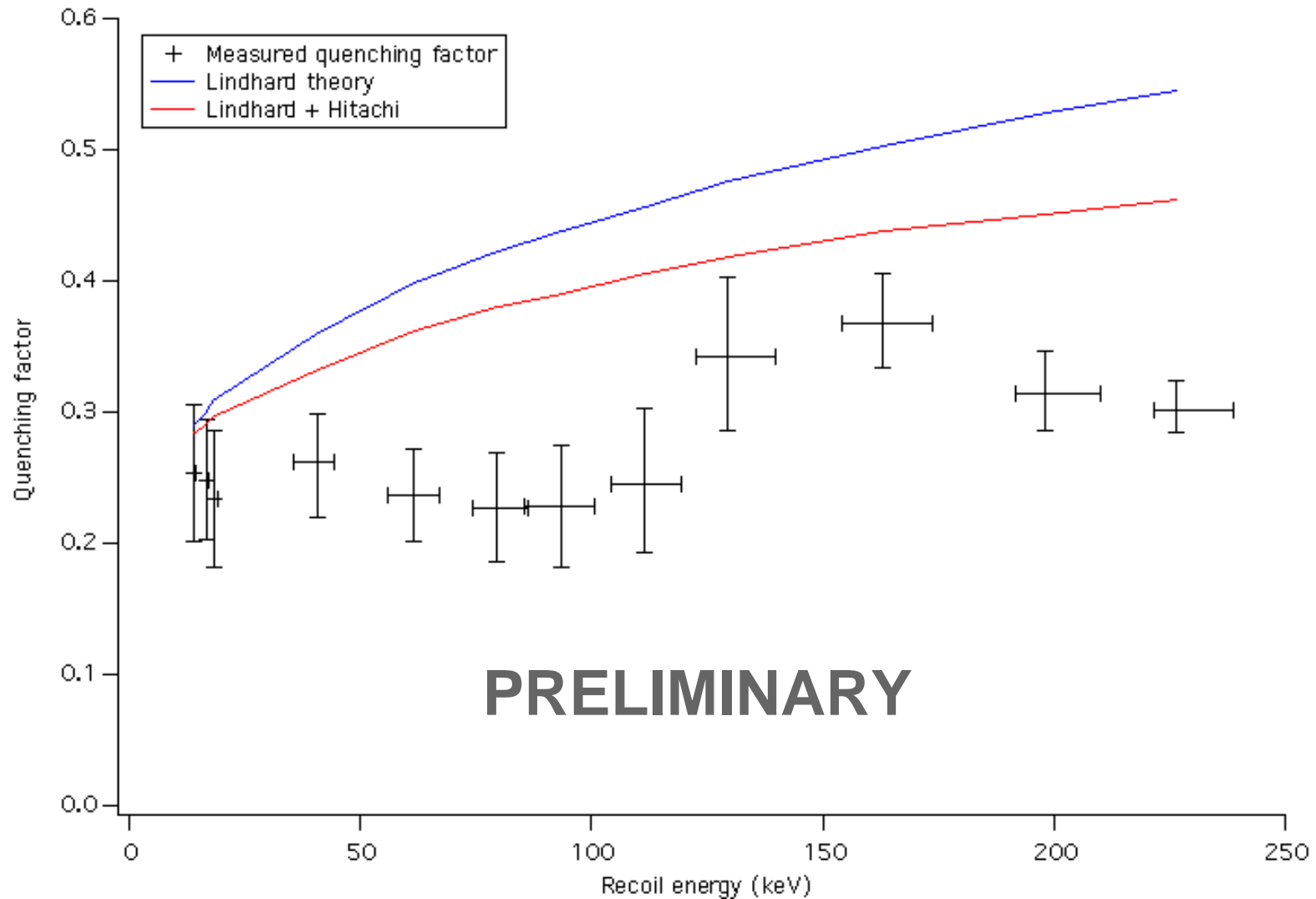
19.0 keV



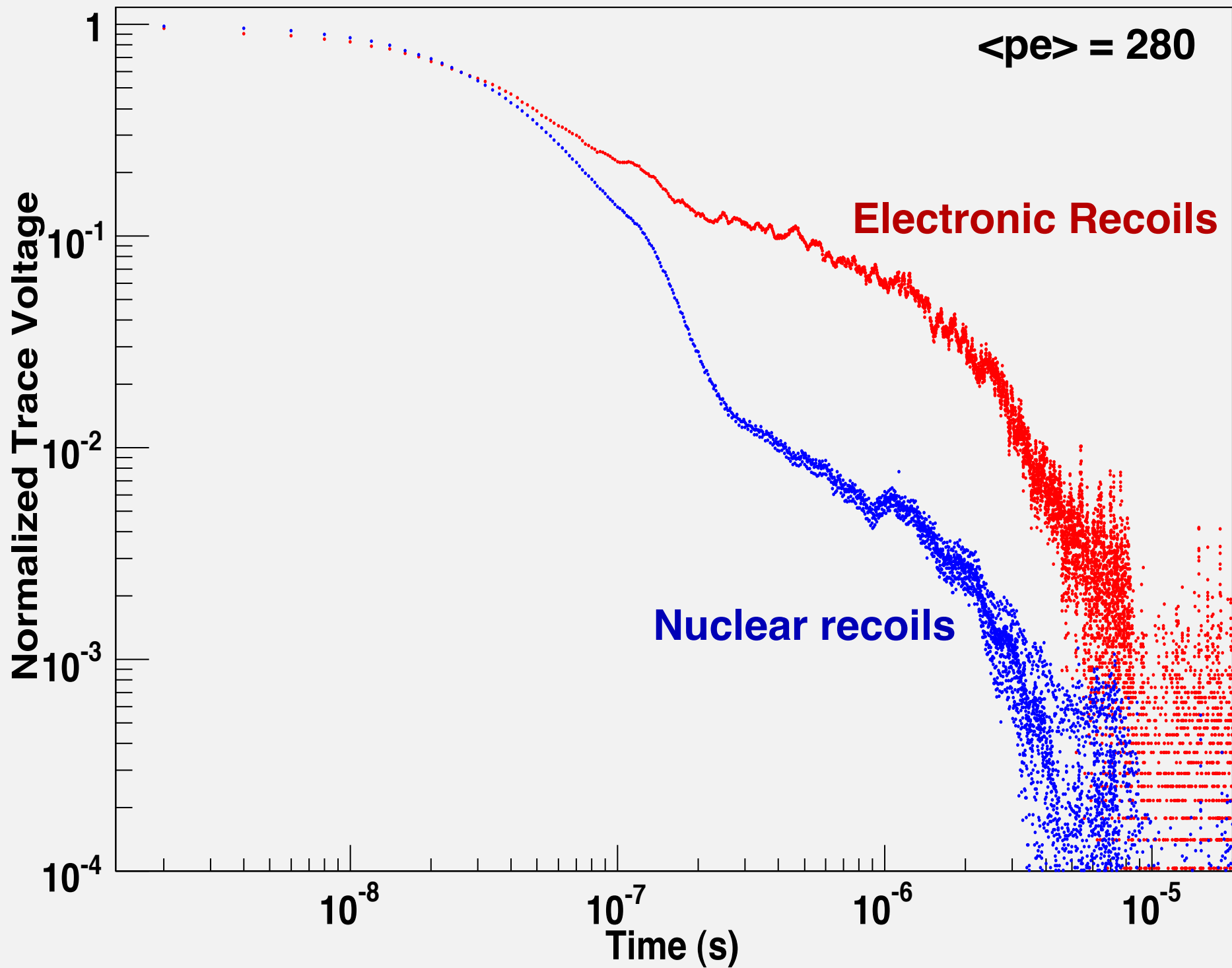
14.5 keV

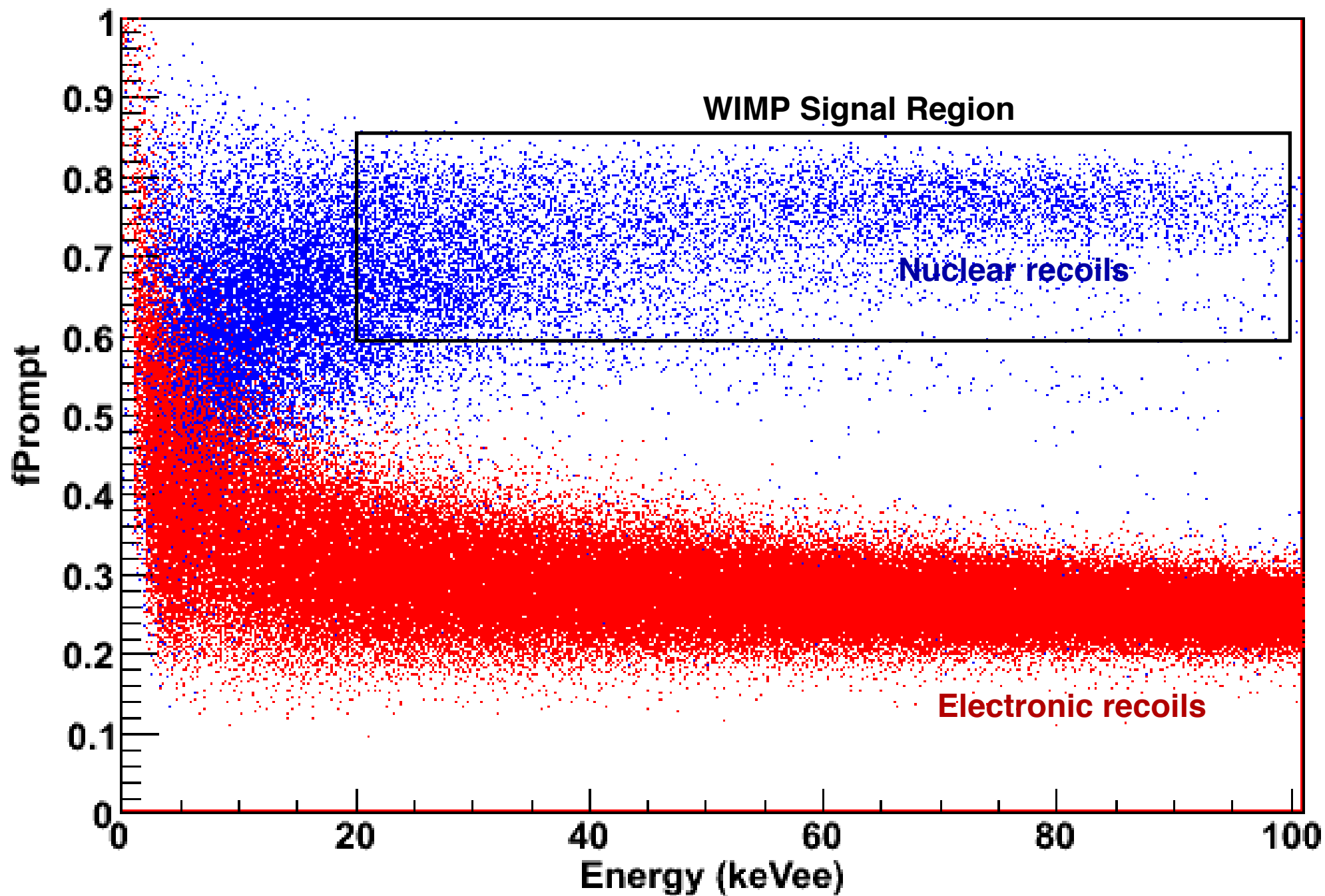


# LAr Quenching Measurements



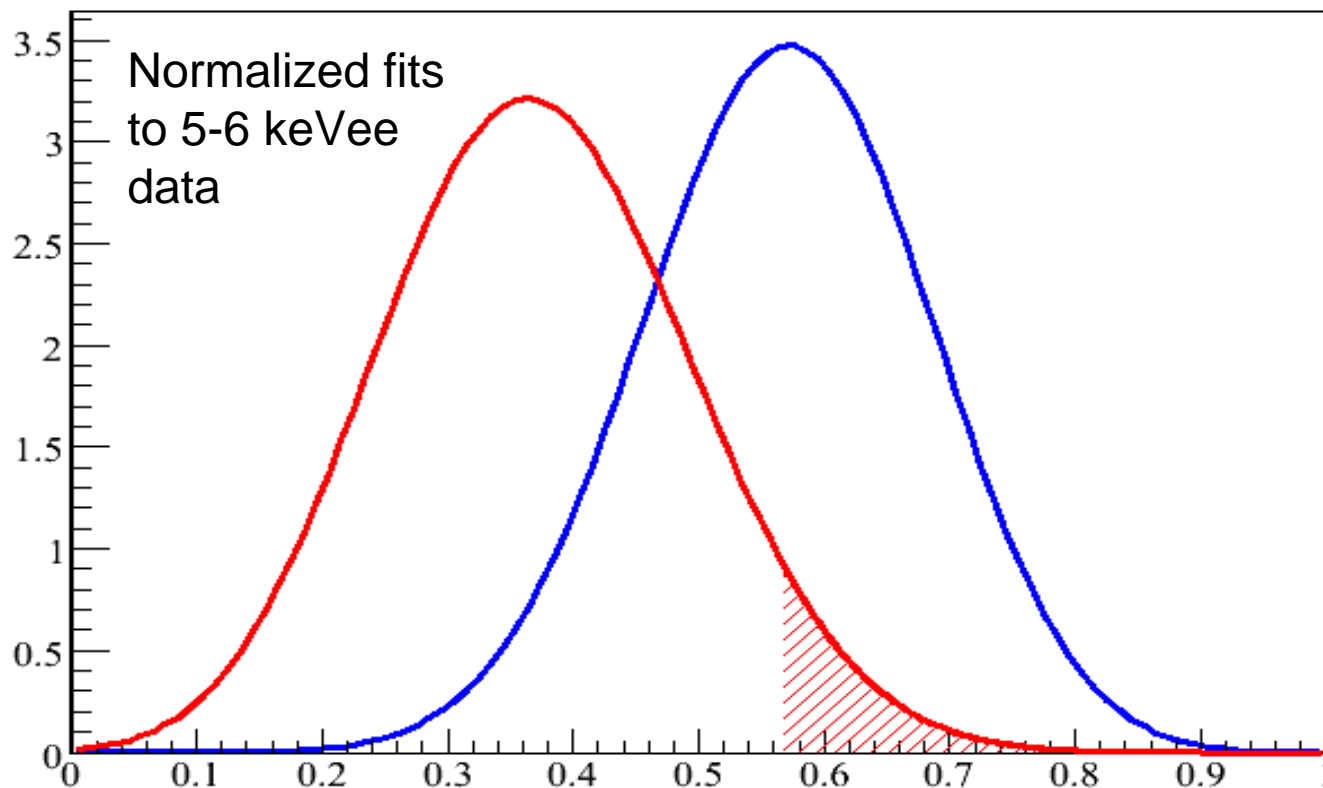
# Time Dependence of Liquid Argon Scintillation



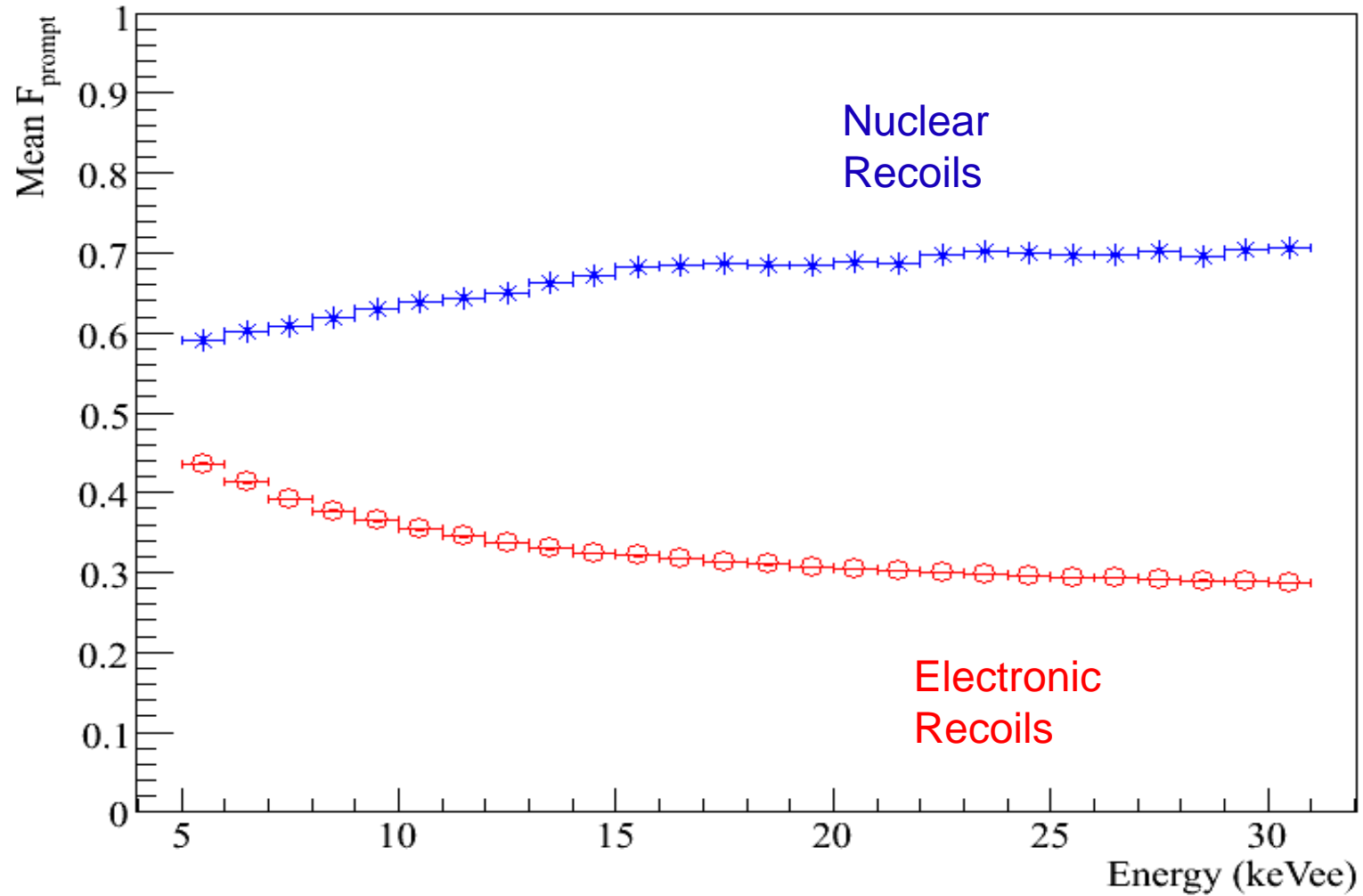


# Analytic Projections

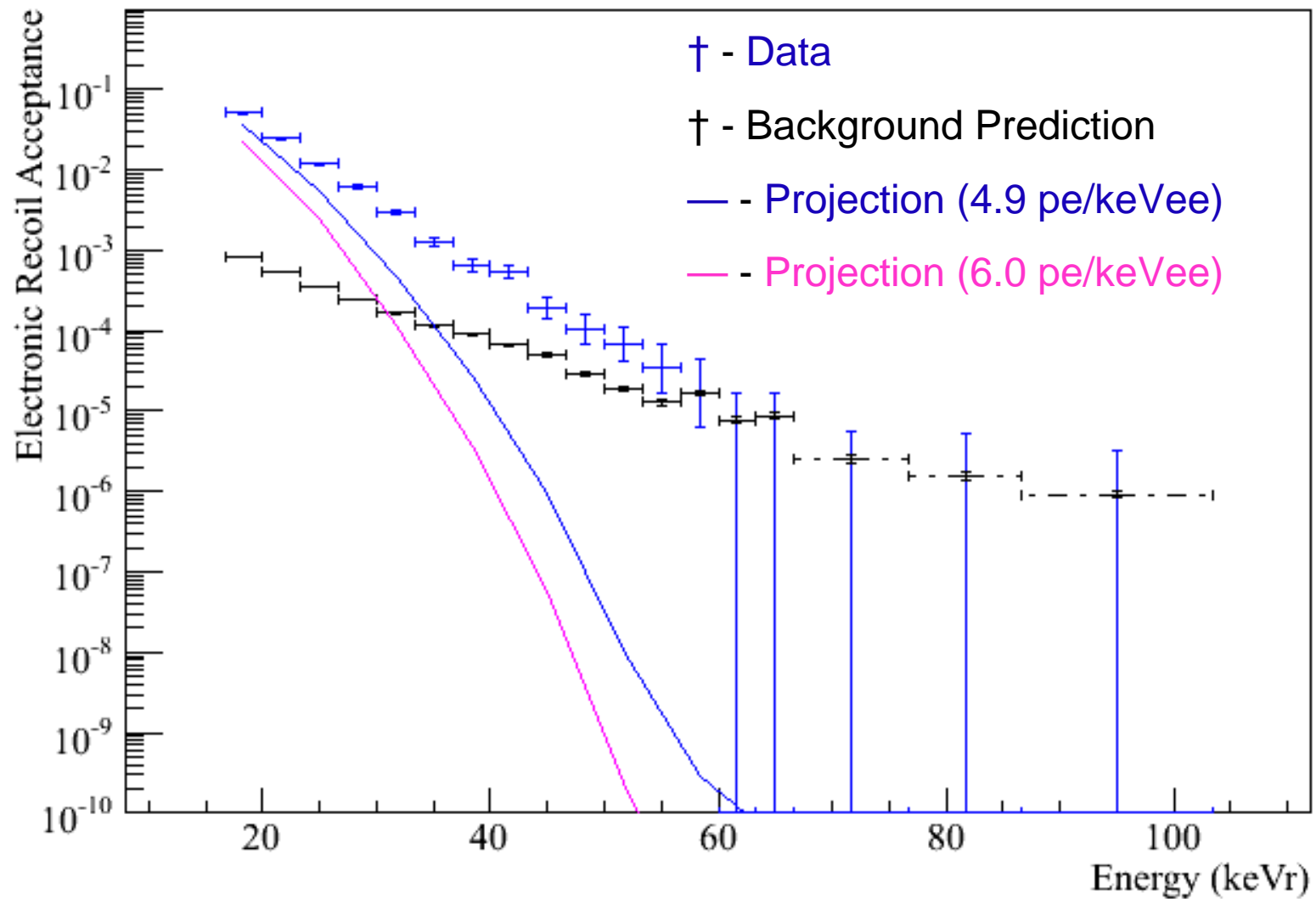
- Analytically integrate area underneath electronic recoil curve above 50% nuclear recoil acceptance level
- Scale fitted binomial widths linearly with light yield, relative to 4.9 pe/keV in microCLEAN



# Mean $F_{\text{prompt}}$ v. Energy

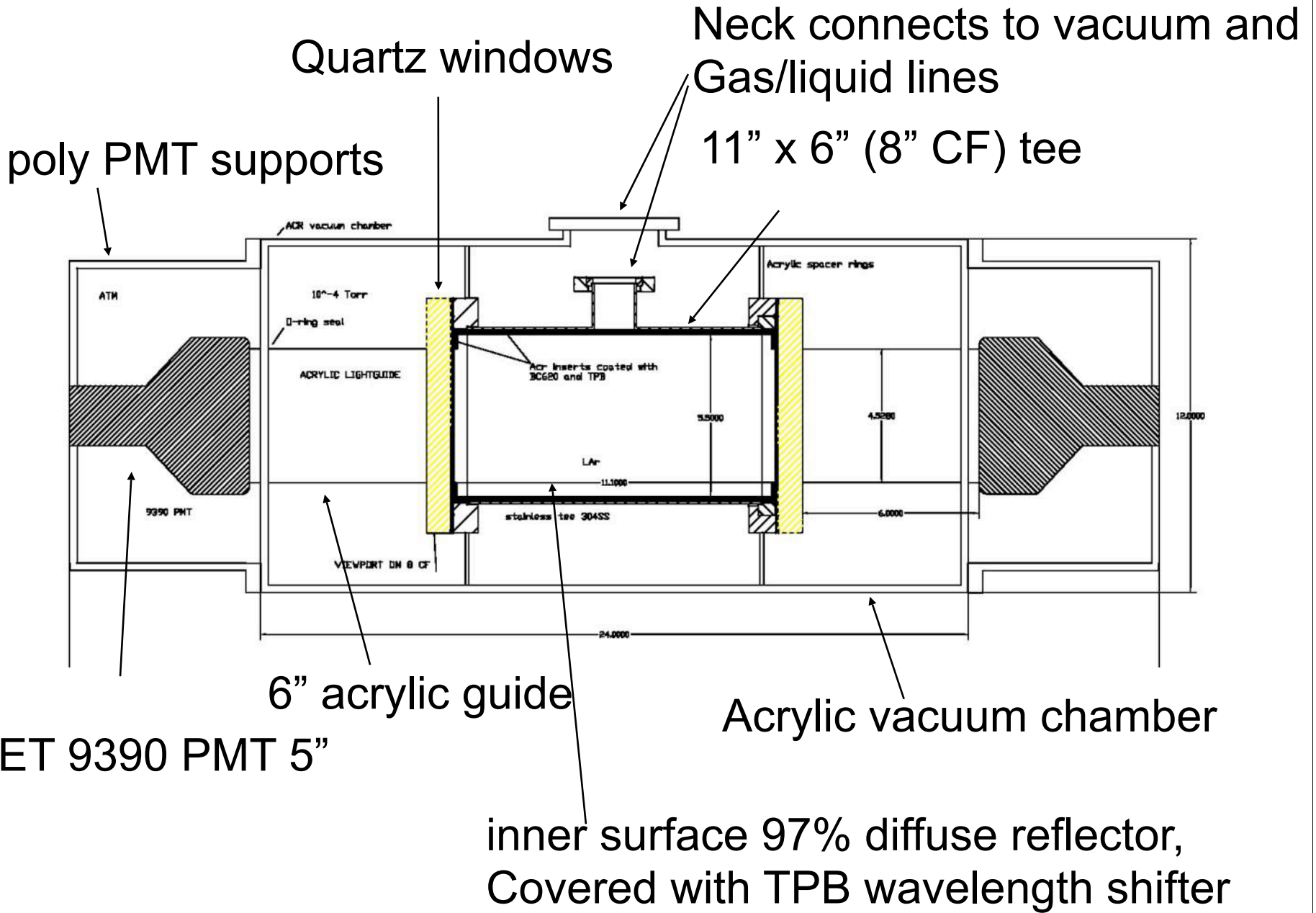






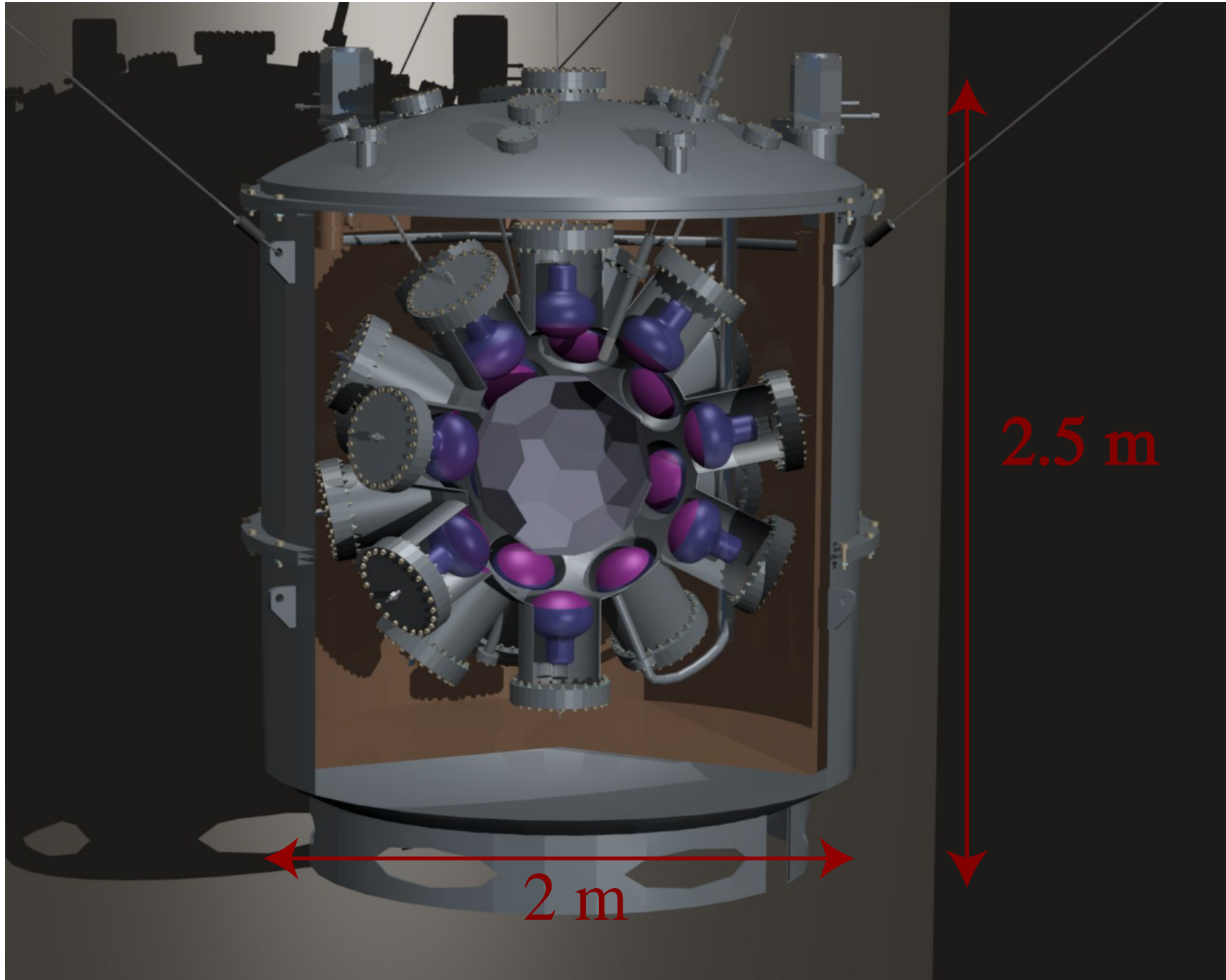
- Binning set to have relatively equal counts in the high energy background bins
- From 60-103 keVr, we have 0 “nuclear recoil events” – better than  $1.25 \times 10^{-6}$  rejection

# DEAP-1 design

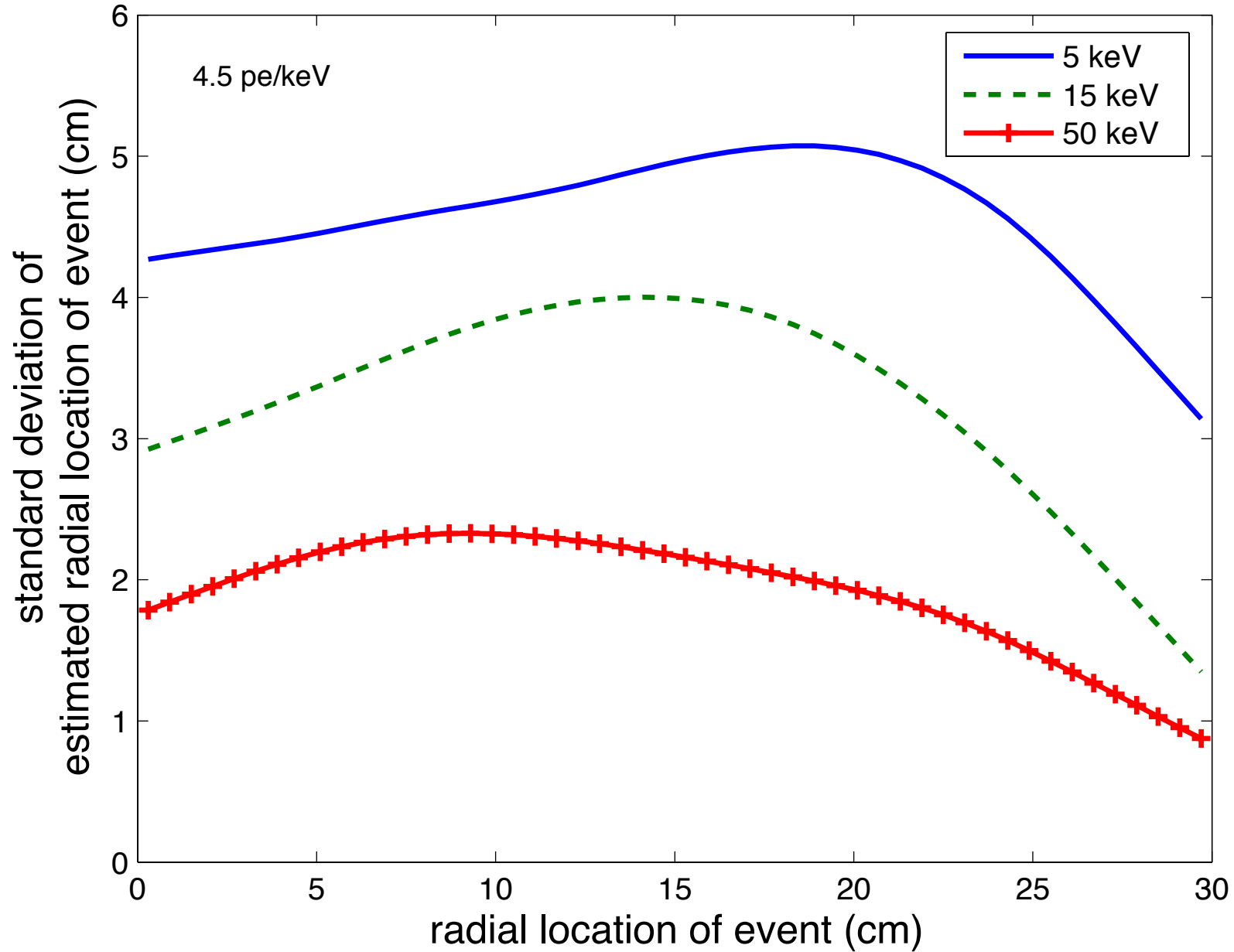


# Mini-CLEAN

Fiducial (total) mass:  $\approx$  100 (~ 500) kg of LAr or LNe. Expected signal yield  $> 6$  pe/keV



# Projected position resolution in mini-CLEAN (Monte Carlo + Maximum Likelihood)



# Neutron Background Study for LAr-Mini-CLEAN

Mei & Hime

<b>Component</b>	<b>Material</b>	<b>U/Th</b>	<b>Yield (n / yr)</b>	<b>Yield in Target (n / kg /yr)</b>	<b>Yield in ROI (n / kg /yr)</b>	<b>Yield in ROI* (n / kg /yr)</b>
<b>Fiducial Sphere</b>	<b>15 kg Quartz</b>	<b>3 ppb</b>	<b>19</b>	<b>0.090</b>	<b>0.047</b>	<b>0.0042</b>
	<b>5 kg SS</b>	<b>3 ppb</b>	<b>4</b>	<b>0.021</b>	<b>0.005</b>	<b>0.0006</b>
<b>PMT Sphere</b>	<b>20 kg SiO<sub>2</sub></b>	<b>30 ppb</b>	<b>256</b>	<b>0.020</b>	<b>0.011</b>	<b>0.0023</b>
	<b>4 kg B<sub>2</sub>O<sub>3</sub></b>	<b>30 ppb</b>	<b>2304</b>	<b>0.340</b>	<b>0.100</b>	<b>0.0238</b>
	<b>85 kg SS</b>	<b>3 ppb</b>	<b>68</b>	<b>0.010</b>	<b>0.003</b>	<b>0.0006</b>
<b>Outer Cryostat</b>	<b>125 kg SS</b>	<b>3 ppb</b>	<b>100</b>	<b>0.020</b>	<b>0.003</b>	<b>0.0008</b>
<b>Total</b>			<b>2751</b>	<b>0.510</b>	<b>0.169</b>	<b>0.032</b>

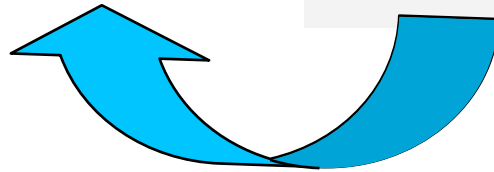
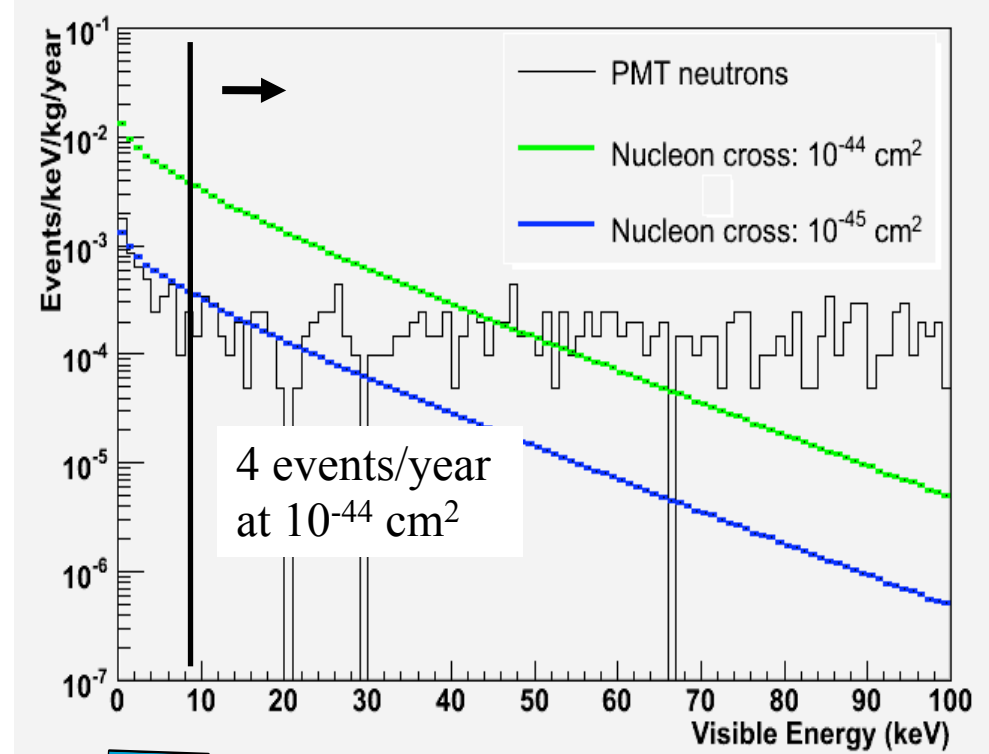
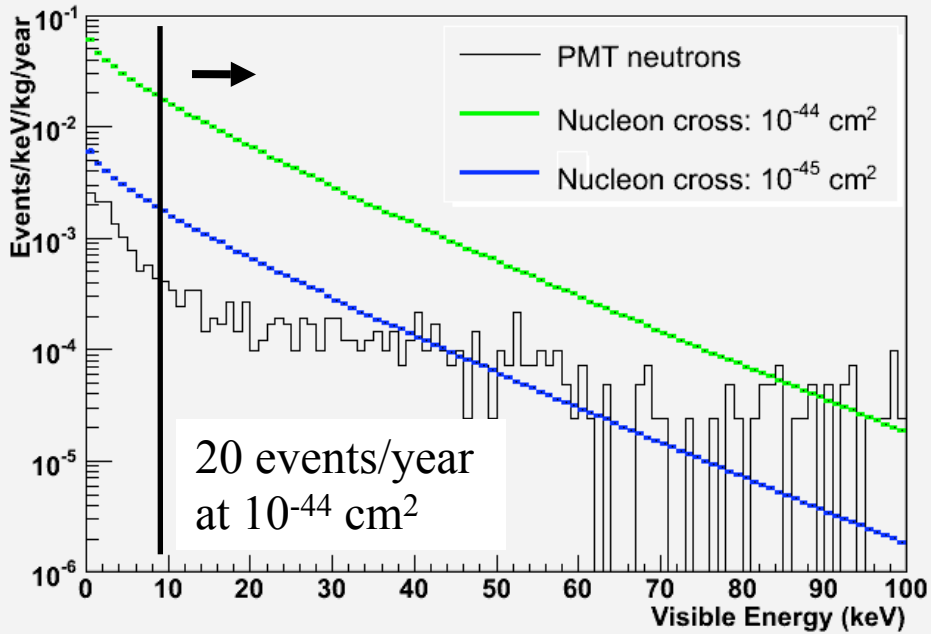
$R_f \sim 26$  cm for  $M_f \sim 100$  kg

$R_{PMT} \sim 38$  cm

$R_C \sim 46$  cm

# 100kg Mini-CLEAN-LAr

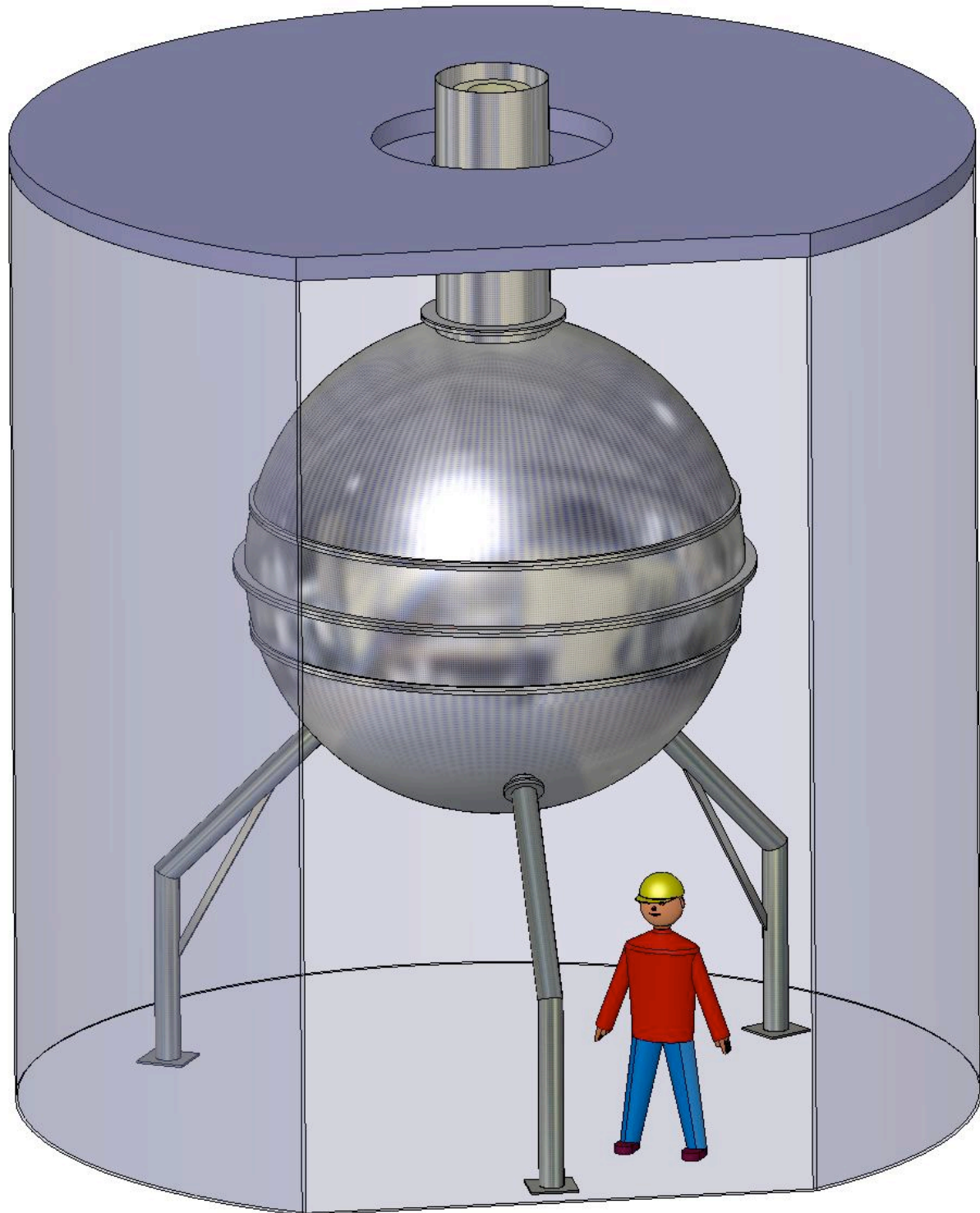
# 100kg Mini-CLEAN-LNe



Exchange Target in Identical Detector to “Exercise” Signal v.s. BGND

assuming a 100 GeV WIMP  $\left\{ \begin{array}{l} S_{\text{Ar}} \sim 5 \times S_{\text{Ne}} \\ B_{\text{Ar}} \sim B_{\text{Ne}} \end{array} \right\}$  & similar energy threshold

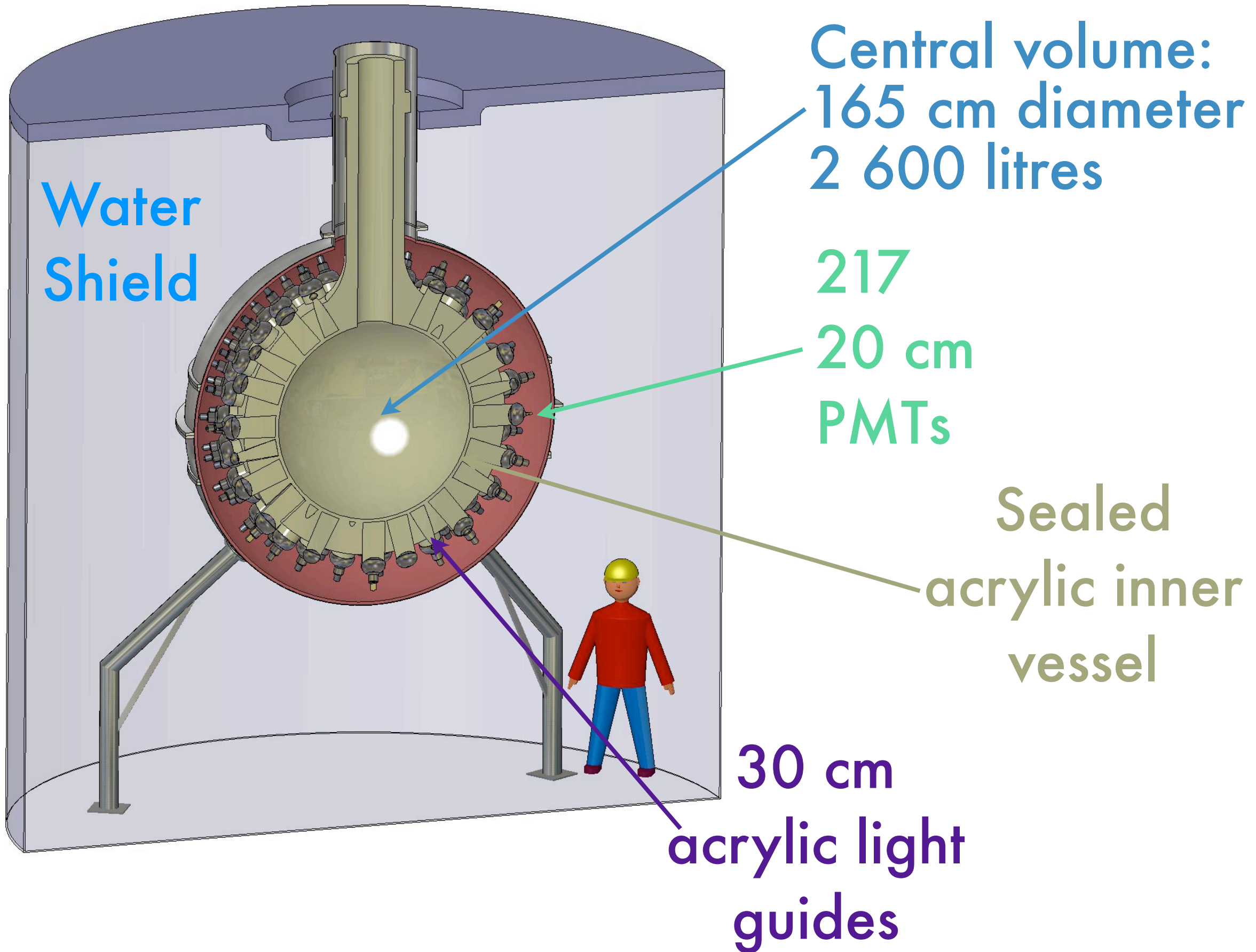
# DEAP/CLEAN 3600



- Dark matter detector with a sensitivity of  $3 \times 10^{-46} \text{ cm}^2$  at 100 GeV
- Liquid neon and liquid argon capable
- 3.1 tonne LNe
- 3.6 tonne LAr



# Cryostat option





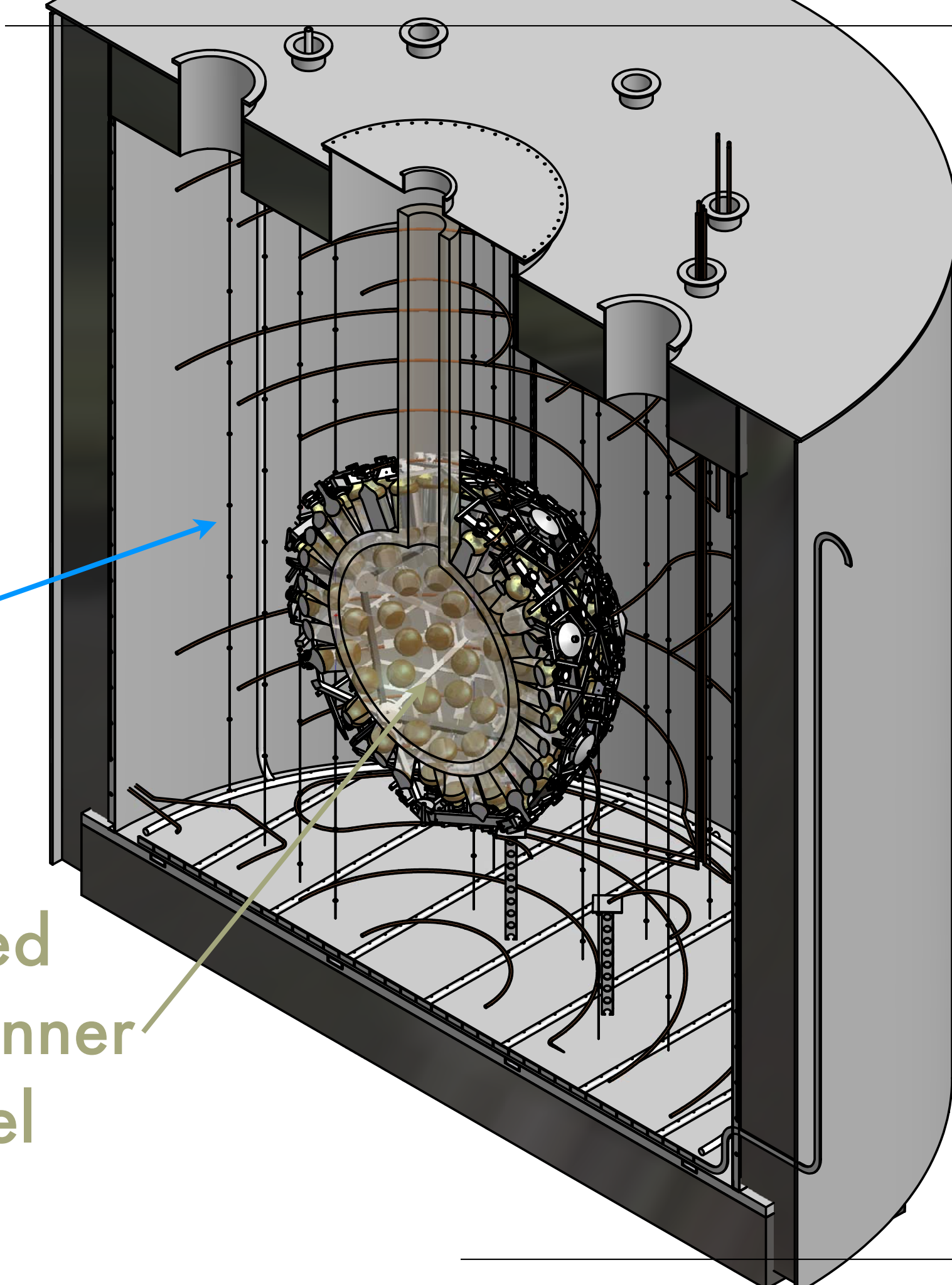
Ice option

Central volume:  
165 cm diameter  
2 600 litres

Ice with no  
cryostat

Reduced  
background  
by eliminating  
steel  
cryostat

Sealed  
acrylic inner  
vessel



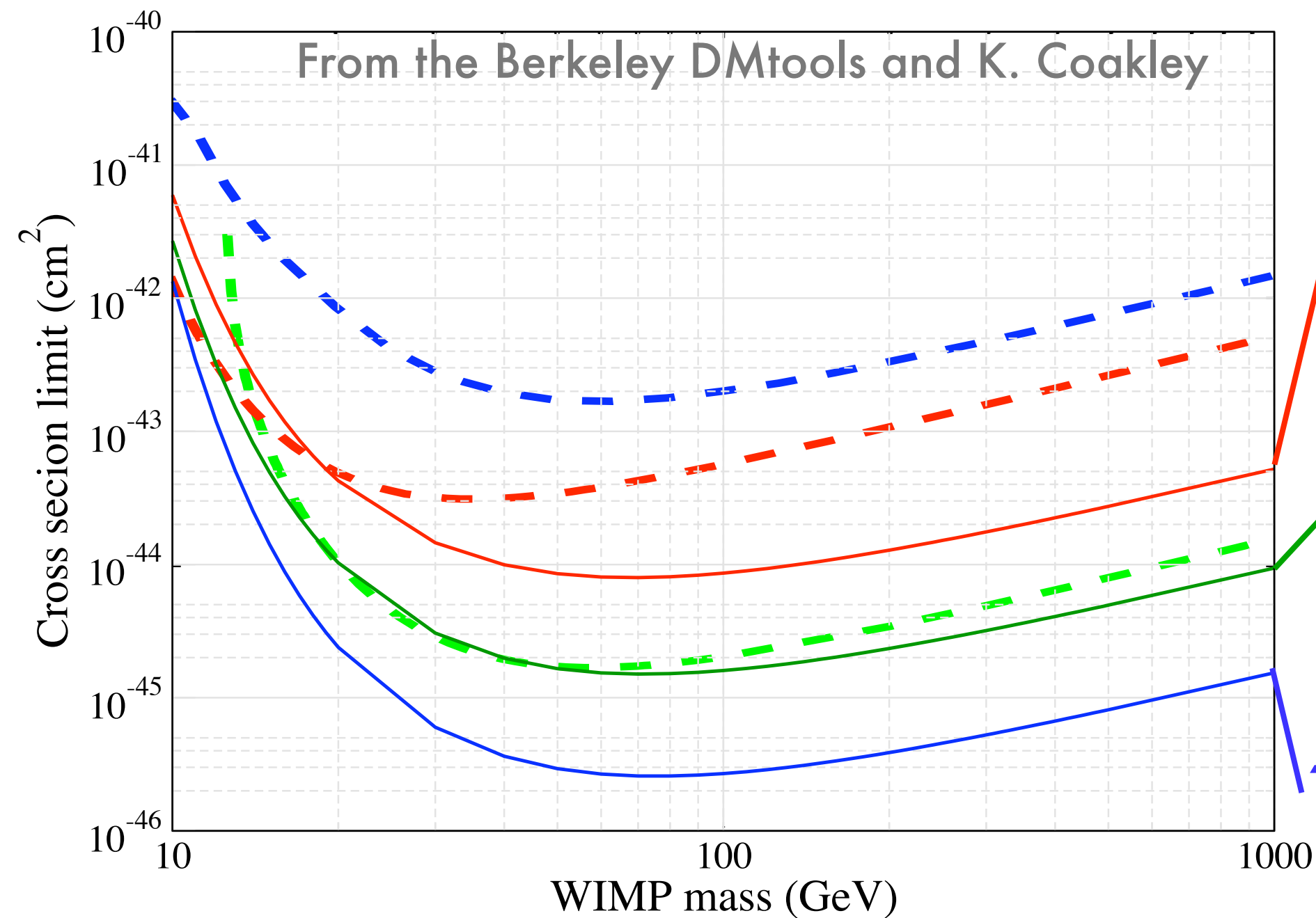
- DATA listed top to bottom on plot
- — — CDMS (Soudan) 2004 + 2005 Ge (7 keV threshold)
  - — — XENON10 2007 (Net 136 kg-d, BG Subtract)
  - — — LUX 300 kg LXe Projection

**Argon:**

**100 Kg Total  
25 Kg Fiducial**

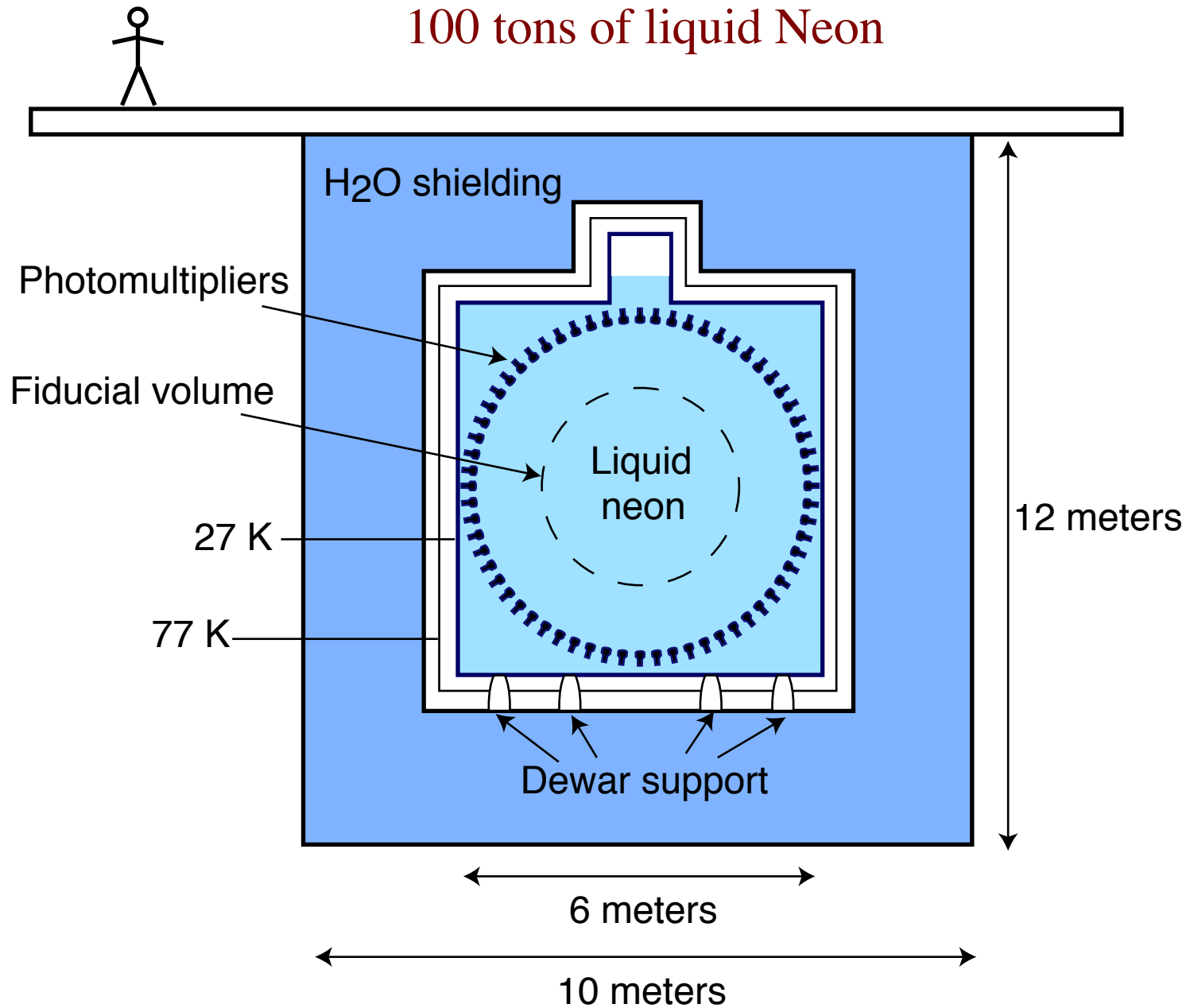
**600 Kg Total  
150 Kg Fiducial**

**4000 Kg Total  
1000 Kg Fiducial**



# CLEAN

100 tons of liquid Neon



Simulations performed with GEANT4

15,000 photons/MeV

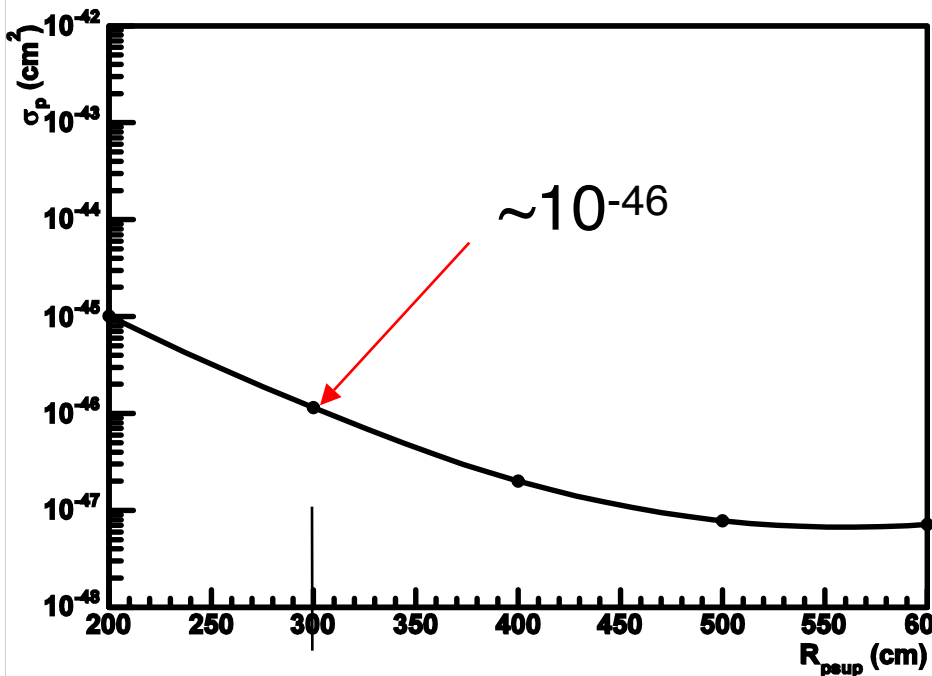
60 cm Rayleigh scattering length

75% PMT coverage, 15% QE

100% photon-to-photon wavelength shifter efficiency

Background dominated by (commercially available) PMT glass: 30 ppb U/Th, 60 ppm K

WIMP sensitivity



300 cm

pp statistical uncertainty

