

Holography and the speed of sound

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Holography and hydrodynamics

- Experimental motivation: Relativistic Heavy Ion Collisions (v_2)
- Traditionally: kinetic description \Rightarrow hydrodynamics
- Discovery of sQGP: hydrodynamics but no kinetic description
 - i.e QFT \Rightarrow hydrodynamics.
- Strong coupling regime of some susy gauge theories can be studied using AdS/CFT (holographic) correspondence.
 - i.e., instead of QFT \Rightarrow kinetic description (Boltzmann) \Rightarrow hydrodynamics,
QFT \Rightarrow holographic description \Rightarrow hydrodynamics
- Kinetic coefficients computed in $\mathcal{N} = 4$ sYM: η , also $\tau, \kappa, \lambda_1, \lambda_2, \lambda_3$ – 2nd order.
- Bulk visc. ζ in some non-conformal susy theories.

Speed of sound

$$c_s^2 = \frac{dp}{d\epsilon}$$

Plays an important role in hydrodynamics.

- Example: Bjorken boost-inv. expansion: $T \sim \tau^{-c_s^2}$.
- Mach cone angle
- $(c_s^2 - 1/3)$ – a measure of the non-conformality of a relativistic system.
- As $T \rightarrow \infty$, $p \sim T^4$ by dimension, i.e., $\epsilon = T dp/dT - p \rightarrow 3p$ and $c_s^2 \rightarrow 1/3$.
- Question: Does c_s^2 always approach $1/3$ from [below](#)?
- Examples: QCD ($N_f = 0$)

$$c_s^2 \simeq \frac{1}{3} + \frac{5N_c}{36\pi} \beta(\alpha) < \frac{1}{3}$$

follows from asymptotic freedom: $\beta < 0$.

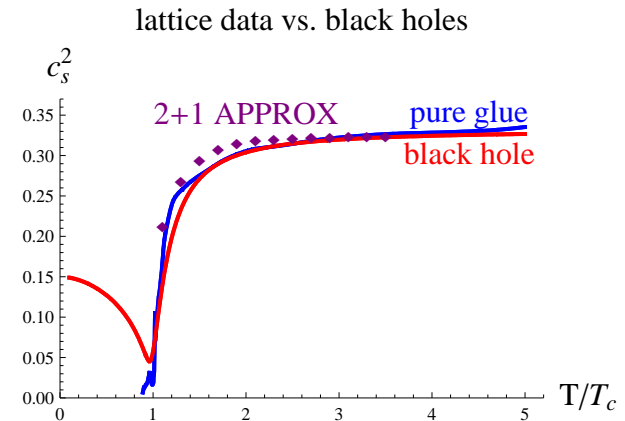
- Strongly interacting systems?

More $c_s^2 < 1/3$ examples

- $N = 2^*$ susy Yang-Mills plasma (Benincasa, Buchel, Starinets)

$$c_s = \frac{1}{\sqrt{3}} \left(1 - \frac{[\Gamma(\frac{3}{4})]^4}{3\pi^4} \left(\frac{m_f}{T}\right)^2 - \frac{1}{18\pi^4} \left(\frac{m_b}{T}\right)^4 + \dots \right).$$

- Lattice QCD (Karsch et al) and holographic mimicking (Gubser-Nellore):



- Not a universal bound!
Counterexample QCD + μ_{isospin} , low T :

$$c_s^2 = \frac{1 - (m_\pi/\mu)^4}{1 + 3(m_\pi/\mu)^4} \rightarrow 1.$$

$$(m_\pi = 0, n \sim \mu^1)$$

- How universal is the result $c_s^2 < 1/3$? What are the prerequisites?

Non-conformality

- Introduce $\theta = \epsilon - 3p$ (as measure of non-conformality)
and $w = \epsilon + p$ (heat function, proxy to T , $dw/dT = c_v + s > 0$).

Express

$$\epsilon = (3w + \theta)/4; \quad p = (w - \theta)/4.$$

$$c_s^2 = \frac{dp}{d\epsilon} = \frac{1 - d\theta/dw}{3 + d\theta/dw}.$$

Then

$$c_s^2 < 1/3 \quad \Leftrightarrow \quad d\theta/dw > 0$$

- Compare/contrast

- Bjorken argues: $p/T^4 \sim \#\text{d.o.f.}$, thus

$$\frac{d}{dT} \left(\frac{p}{T^4} \right) = \frac{\theta}{T^5} > 0$$

- Appelquist-Cohen-Schmaltz-Shrock:

$$\left[\frac{p}{T^4} \right]_{\text{IR}}^{\text{UV}} = \int_{\text{IR}}^{\text{UV}} \frac{dT}{T^5} \theta > 0$$

Bjorken on $\theta > 0$

Other general features of the hydrodynamic expansion follow from the positivity condition on the trace of the energy-momentum tensor

$$T^\mu{}_\mu \geq 0 \quad (33)$$

which is true under quite general circumstances.⁹

⁹This is far from guaranteed, but is always true when the fluid can be considered a collection of noninteracting quanta.

Holographic model

- Bottom-up approach: Take a “minimal” set of operators: $T^{\mu\nu}$ and a scalar \mathcal{O} on the r.h.s. of scale (trace) anomaly eqn. (in QCD $\theta = \frac{\beta}{8\pi\alpha^2} \langle \text{tr } F^2 \rangle$) to describe conformality violation.

- Holographic duality requires 5d fields: $g_{\mu\nu}$ (dual to $T^{\mu\nu}$) and ϕ (dual to \mathcal{O}).

$$S_5 = \frac{1}{2\kappa^2} \left[\int_M d^5x \sqrt{-g} \left(R - V(\phi) - \frac{1}{2}(\partial\phi)^2 \right) - 2 \int_{\partial M} d^4x \sqrt{-\gamma} K \right],$$

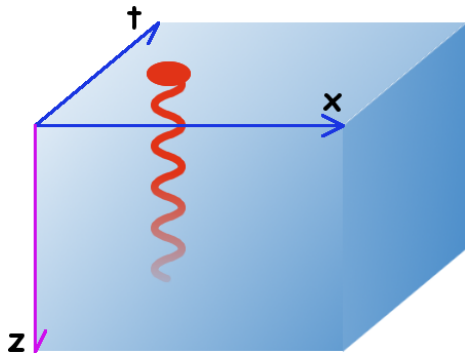
with *generic* scalar potential $V(\phi)$ (tuned by Gubser-Nellore).

- $1/\kappa^2 \sim \# \text{ d.o.f.}$

- Recipe for calculating a correlator of, e.g., $T^{\mu\nu}$:

Vary $z = 0$ boundary value, $g^{\mu\nu}(x, 0)$, then

$$\langle T^{\mu\nu}(x) \rangle \sim \frac{\delta S}{\delta g_{\mu\nu}(x, 0)}.$$



Background solution

● Most general (up to coordinate transforms):

$$ds^2 = \frac{1}{z^2} \left(-f(z) dt^2 + d\vec{x}^2 \right) + e^{2B(z)} \frac{dz^2}{z^2 f(z)}.$$

where f , B , and ϕ obey

$$\dot{B} = -\frac{1}{6} \dot{\phi}^2,$$

$$\ddot{f} = (4 + \dot{B}) \dot{f},$$

$$-6\dot{f} + f(24 - \dot{\phi}^2) + 2e^{2B} V(\phi) = 0,$$

$$\ddot{\phi} f + \dot{\phi} (\dot{f} - f(4 + \dot{B})) - e^{2B} V'(\phi) = 0,$$

where a dot denotes a $\log z$ derivative, e.g., $\dot{\phi} = z d\phi/dz$, while $V' = dV/d\phi$.

● The only boundary conditions at $z = \varepsilon \rightarrow 0$:

$$f(\varepsilon) = 1 \quad \text{and} \quad \phi(\varepsilon) = c \varepsilon^{\Delta_-}.$$

where c is the source for \mathcal{O} , i.e., $[c] = \Delta_-$; $[\mathcal{O}] = \Delta_+$; $\Delta_+ + \Delta_- = 4$.

Background solution

- Equation $\ddot{f} = (4 + \dot{B}) \dot{f}$ can be integrated once:

$$\dot{f} = -wz^4 e^B.$$

integration constant $w > 0$ to have a horizon $f(z_H) = 0$.

- Eq. $\ddot{\phi} f + \dots$ linearizes near $z = 0$ and is solved by (with $\phi(\varepsilon) = c\varepsilon^{\Delta_-}$)

$$\phi(z) \rightarrow (c - d\varepsilon^{\Delta_+ - \Delta_-}) z^{\Delta_-} + d z^{\Delta_+} + \dots,$$

where $\Delta_{\pm} = 2 \pm \sqrt{4 + m^2}$, with $m^2 \equiv V''(0)$.

d is fixed by condition $|\phi(z_H)| < \infty$.

Substituting this solution back into S_5 one finds (Klebanov-Witten):

$$\langle \mathcal{O} \rangle = \partial S_5 / \partial c = -d(\Delta_+ - \Delta_-).$$

- Two-parameter (w, c) family of solutions.

Scaling $z \rightarrow z/\lambda$, $c \rightarrow c\lambda^{\Delta_-}$ and $w \rightarrow w\lambda^4$ relates different solutions.

One-point functions

- Consider variable boundary condition on the metric

$$g_{\mu\nu}(\varepsilon) = h_{\mu\nu} \varepsilon^{-2}.$$

where $h_{\mu\nu}$ is the metric in 4D theory (in general, curved).

- Holographic correspondence then gives us

$$\langle T^{\mu\nu} \rangle = -2 \frac{\delta S_5}{\delta h_{\mu\nu}} \quad \Rightarrow \quad \langle T^{tt} \rangle = -\frac{6}{\varepsilon^4} e^{-B(\varepsilon)}; \quad \langle T^{xx} \rangle = w - \langle T^{tt} \rangle.$$

and (subtracting T -indep. constant) $(B(\varepsilon) \leftarrow -6f + f(24 - \dot{\phi}^2) + 2e^{2B} V(\phi) = 0)$

$$\epsilon = \frac{3}{4} w - \frac{1}{4} c d \Delta_- (\Delta_+ - \Delta_-) \quad \text{and} \quad p = \frac{1}{4} w + \frac{1}{4} c d \Delta_- (\Delta_+ - \Delta_-).$$

or

$$\theta = \epsilon - 3p = -c d \Delta_- (\Delta_+ - \Delta_-) = c \langle \mathcal{O} \rangle \Delta_-,$$

- Recognizable and derivable directly on the field-theory side.

Speed of sound

● Recall

$$c_s^2 = \frac{dp}{d\epsilon} = \frac{1 - d\theta/dw}{3 + d\theta/dw}.$$

with $\theta = \Delta_- c \langle \mathcal{O} \rangle$ and $\langle \mathcal{O} \rangle = -d(\Delta_+ - \Delta_-)$, and no approximation,

$$c_s^2 = \frac{dp}{d\epsilon} = \frac{1 + c \Delta_- (\Delta_+ - \Delta_-) (\partial d / \partial w)}{3 - c \Delta_- (\Delta_+ - \Delta_-) (\partial d / \partial w)}.$$

High temperatures

At large w (i.e., high T), z_h is small and $B \approx 0$:

$$f(z) = 1 - w z^4 / 4.$$

$$\left(1 - \frac{1}{4} w z^4\right) \phi'' - \left(\frac{3}{z} + \frac{w z^4}{4}\right) \phi' - \frac{m^2}{z^2} \phi = 0.$$

$$\begin{aligned} \phi(z) = & c z^{\Delta_-} {}_2F_1\left(\Delta_-/4, \Delta_-/4, \Delta_-/2, w z^4/4\right) \\ & + d z^{\Delta_+} {}_2F_1\left(\Delta_+/4, \Delta_+/4, \Delta_+/2, w z^4/4\right), \end{aligned}$$

and $|\phi(z_h)| < \infty$ gives

$$d = -c w^{(\Delta_+ - \Delta_-)/4} D(\Delta_-),$$

where the function $D(\Delta_-) = 1/D(\Delta_+)$ is

$$D(\Delta_-) = \frac{\pi 2^{\Delta_-}}{2 - \Delta_-} \cot(\pi \Delta_- / 4) \frac{\Gamma(\Delta_- / 2)^2}{\Gamma(\Delta_- / 4)^4} > 0.$$

High temperatures

- Substituting we find our main result

$$c_s^2 = \frac{1}{3} - \frac{1}{9} c^2 w^{-\Delta_-/2} \Delta_- (\Delta_+ - \Delta_-)^2 D(\Delta_-) + \dots < \frac{1}{3}.$$

universally for all $0 < \Delta < 4$.

- Same result, different approach — Cherman, Cohen, Nellore.
- Dimensionful parameters: c (b.c. on ϕ) and w (integration const).

At high T , dimensionless expansion parameter is c/T^{Δ_-} . High T is small c .

- Origin of the universality:

$z_H \sim w^{-1/4} \rightarrow 0$. Thus ϕ remains small \Rightarrow only $V''(0) = \Delta(\Delta - 4)$ matters.

- Why $c_s^2 < 1/3$?

More general argument

- Recall

$$c_s^2 < 1/3 \quad \Leftrightarrow \quad d\theta/dw > 0$$

- At high T (small c):

$$\theta = \Delta_- c \langle \mathcal{O} \rangle = \Delta_- c^2 \chi_{\mathcal{O}} + \dots$$

The susceptibility $\chi_{\mathcal{O}} \equiv \partial \langle \mathcal{O} \rangle / \partial c$ can be related to the effective potential for $\langle \mathcal{O} \rangle$, (in holography $W = -S_5$)

$$\Gamma(\langle \mathcal{O} \rangle) \equiv W(c) + c \langle \mathcal{O} \rangle,$$

as

$$\chi_{\mathcal{O}} = 1/\Gamma''(0).$$

- Stability implies $\Gamma''(0) > 0$, and thus $\chi_{\mathcal{O}} > 0$. This means $\theta > 0$. (cf. Bjorken)

- For $d\theta/dw > 0$ curvature $\Gamma''(0)$ must decrease with T . Strange? Opposite to ϕ^4 .

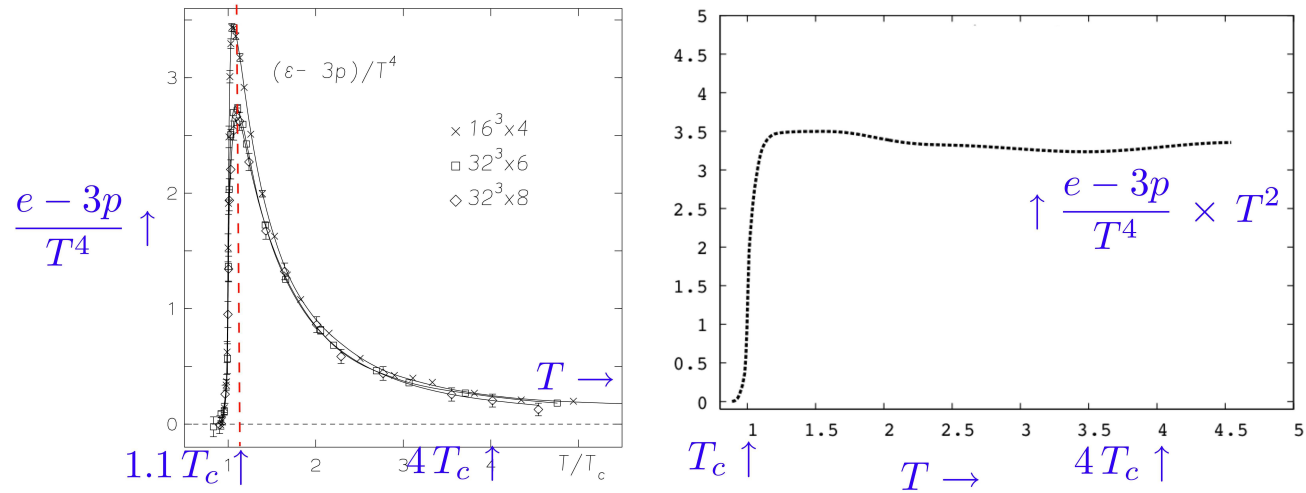
Dimension counting: $[\chi_{\mathcal{O}}] = \Delta_+ - \Delta_- > 0$, which means $\chi_{\mathcal{O}} \sim T^{\Delta_+ - \Delta_-}$ and increases with T .

Conclusions and comments

- In a wide class of strongly interacting theories $c_s^2 \rightarrow 1/3$ from [below](#).
- In holography, the correction is a universal (negative) function of Δ , times c^2 , times $w^{-\Delta-1/2}$ (i.e., $T^{-2\Delta-1}$).
- Can one extend the argument to all T ?
- $\mu \neq 0$?

Comments

- Looking at lattice data on YM Pisarski observes: $\theta \sim w^{1/2}$.



This would take $\Delta_- \approx 1$. Is there a $[\mathcal{O}] \approx 3$ operator?

Happy Birthday, Misha!